

Optimizing x-ray mirror thermal performance using variable-length cooling for high-repetition-rate FELs.

Corey Hardin, Venkat N. Srinivasan, Lope Amores,
Nicholas M. Kelez, Daniel S. Morton, Peter M. Stefan,
Daniele Cocco
August 28th, 2016

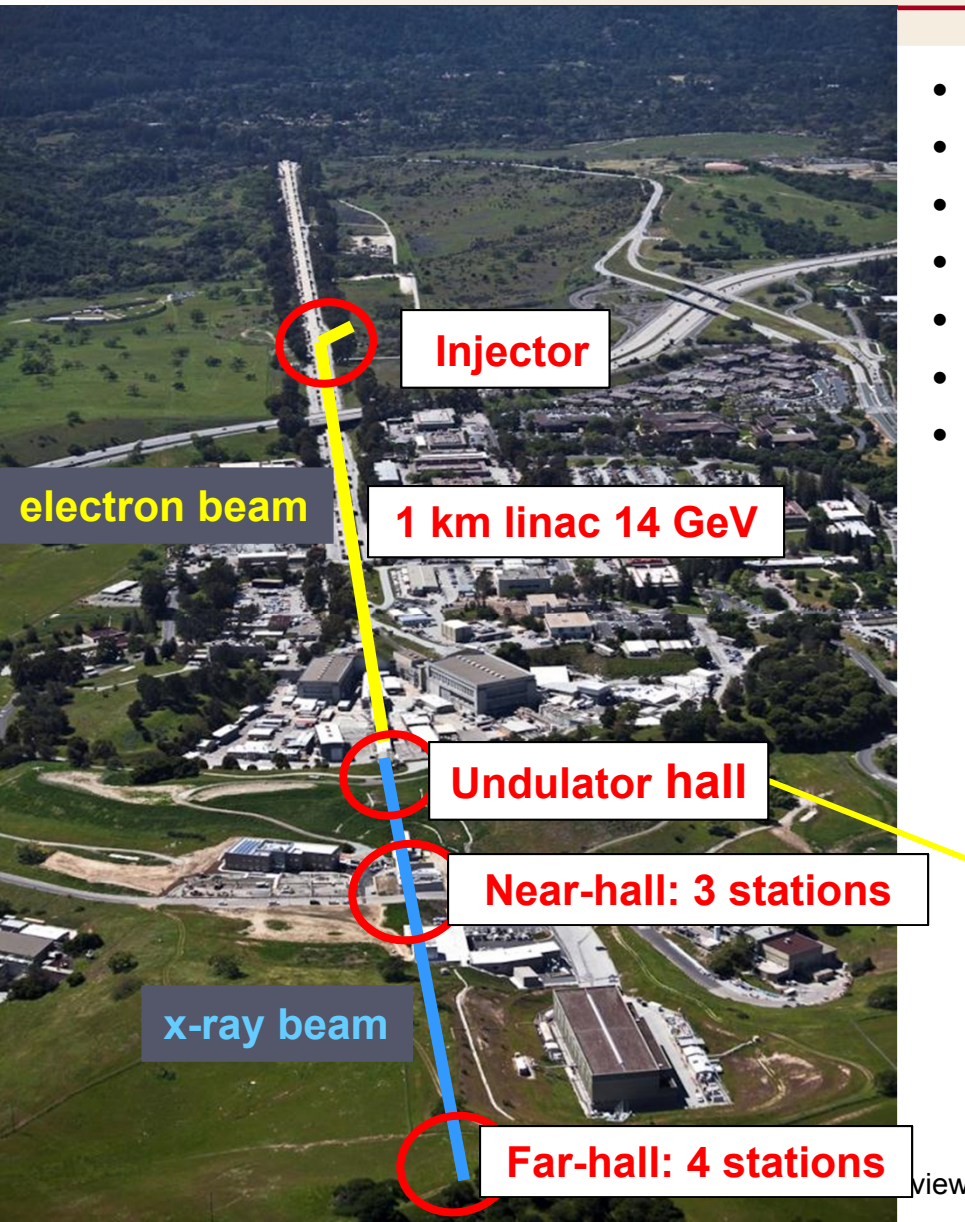
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M.L. Ng, L. Zhang, D. Rich,
M. Rowen, M. Idir, Jtec, Axilon,
D. Flath, A. Wallace, T. Pardini,
L. Christensen, K. Alano, L. Raimondi,
T. Rabedeau, L. Ladao, M. Church,
L. Lee, D. Stefanescu, D. Fritz,
A. Catalano.....

Linac Coherent Light Source

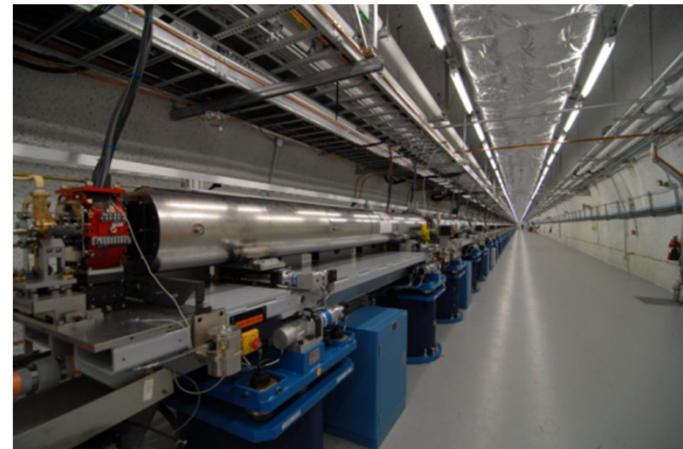


LCLS is part of the SLAC National Accelerator Laboratory located within the Stanford University campus in Menlo Park, CA.

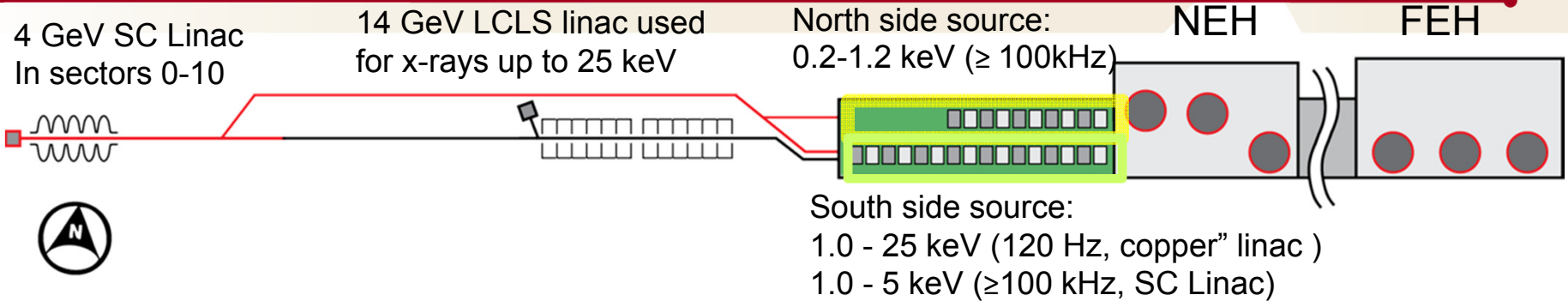
Linac Coherent Light Source



- Uses last 1 km of 3-km linac
- e-beam energy: 2-14 GeV
- X-rays produced by ~100 m undulator
- Photon energy range: 250 eV-10 keV
- Pulse length: 1 -100 fs
- 6 experimental stations
- >600 On Site Users per from over 30 countries



LCLS - II



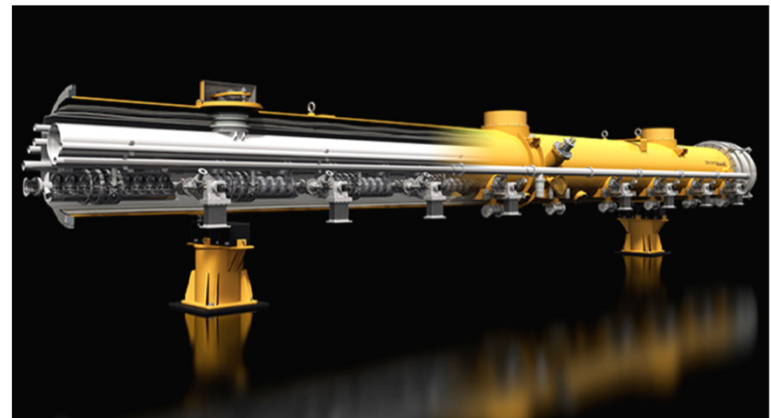
SLAC

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Fermilab

Jefferson Lab



LCLS - II

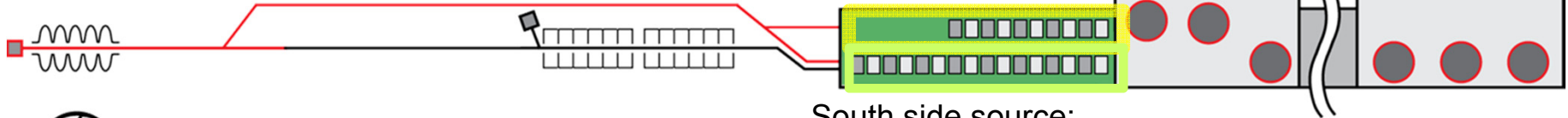
4 GeV SC Linac
In sectors 0-10

14 GeV LCLS linac used
for x-rays up to 25 keV

North side source:
0.2-1.2 keV ($\geq 100\text{kHz}$)

NEH

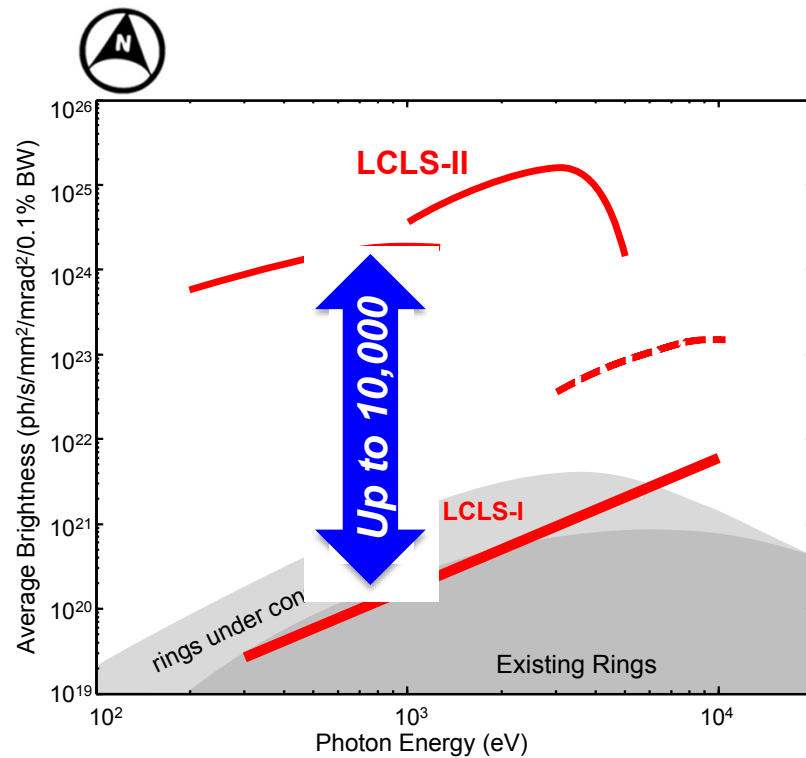
FEH



South side source:

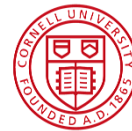
1.0 - 25 keV (120 Hz, copper" linac)

1.0 - 5 keV ($\geq 100\text{ kHz}$, SC Linac)



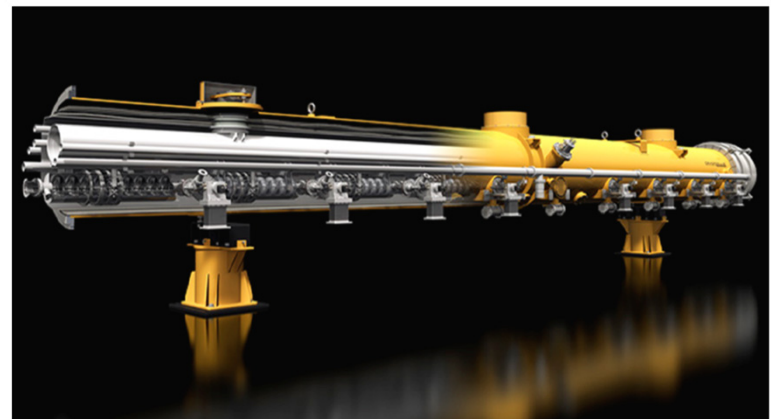
SLAC

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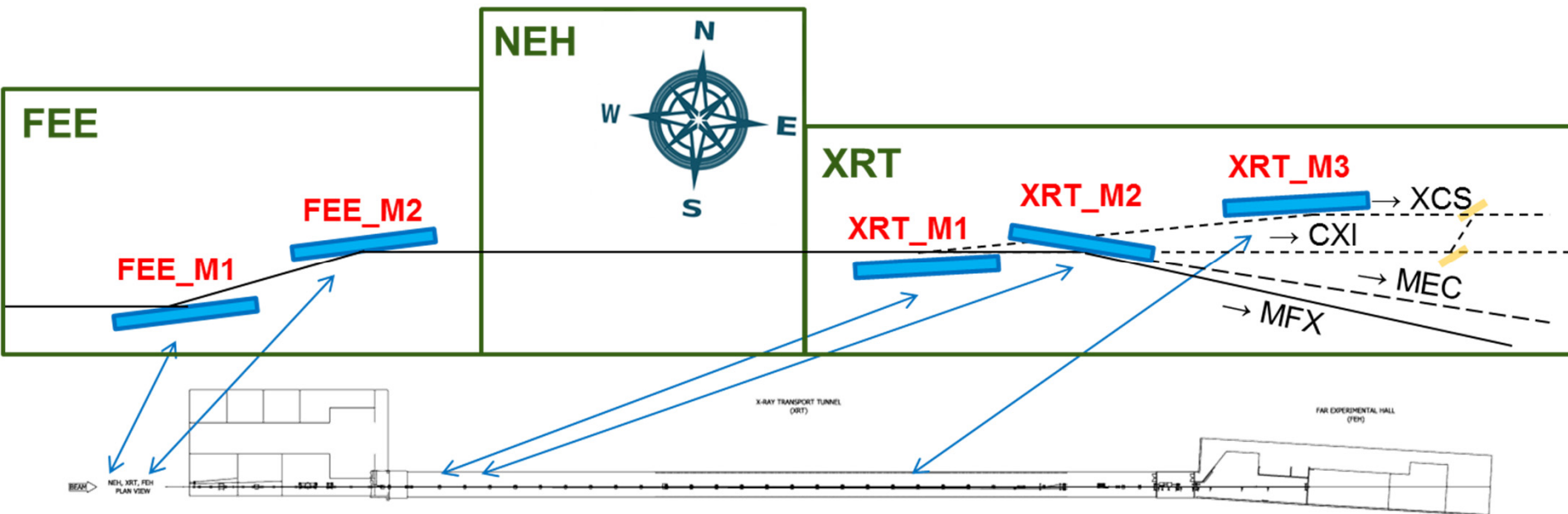
Fermilab

Jefferson Lab



**From the current 2-300 mW to
up to 600 W on the optics**

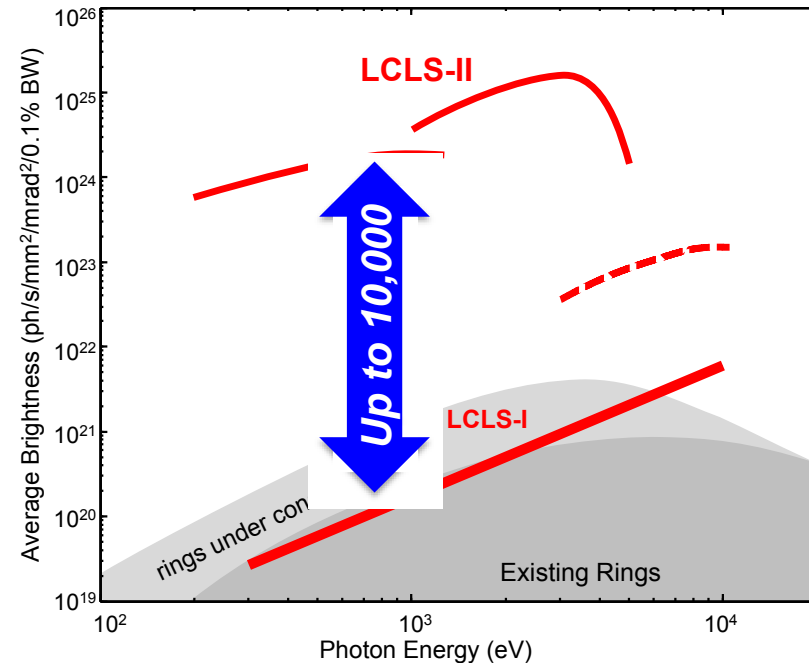
LCLS-I Flat Mirror System Locations



Ack: D. Cocco

LCLS Flat Mirror System Upgrade

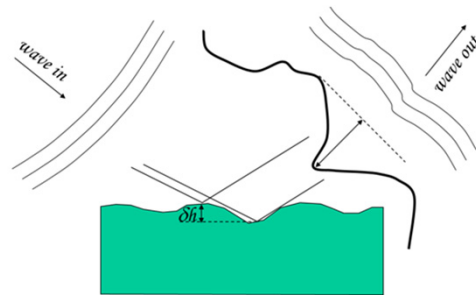
- We need to upgrade the flat mirrors at LCLS to accommodate LCLS II upgrade
- Mirror Requirements:
 - Preserve wavefront in and out of focus
 - Optimized for lower power but with good performance at higher power (200 W)
 - Accommodate Variable beam footprint (from 70 to 400 mm FWHM)
 - Limit induced vibration
 - Avoid LN2, to limit contamination absorption



From the current 2-300 mW to up to 600 W on the optics

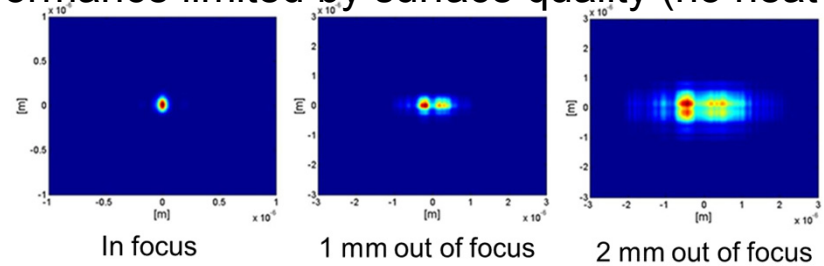
Motivation - Mirror shape errors requirements

- Preserve wavefront in and out of focus
- Optimized for lower power but with good performance at higher power (200 W)



From what we have now (1-2 nm rms shape error mirrors, limited acceptance)

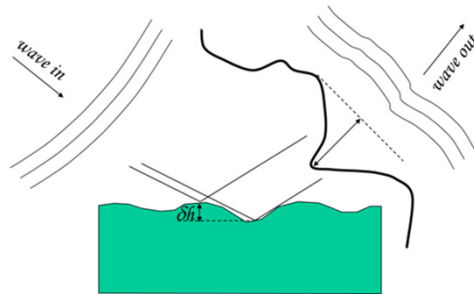
Performance limited by surface quality (no heat load)



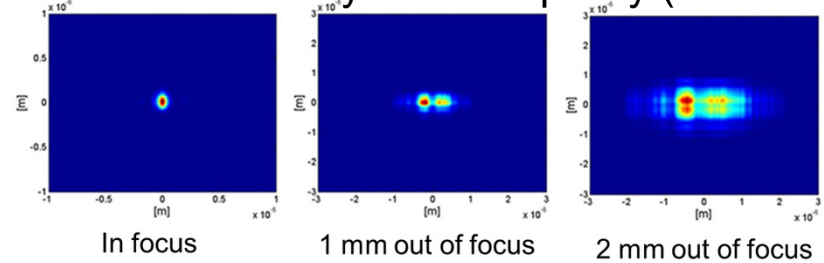
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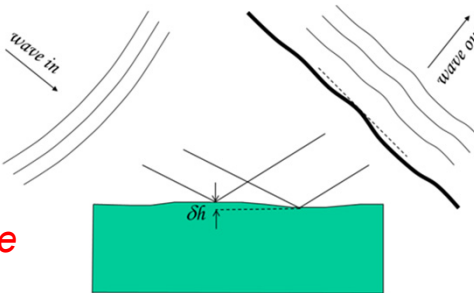
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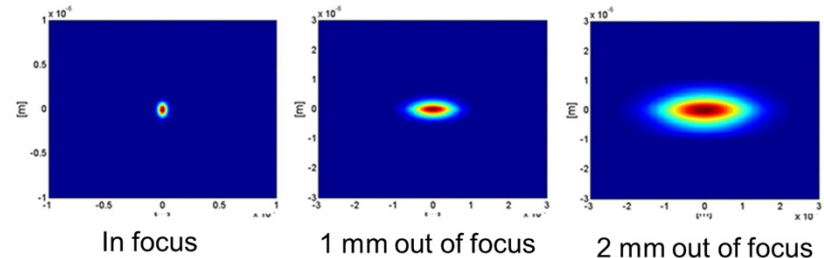
Performance limited by surface quality (no heat load)



..to what we need to have (< 0.5 nm rms, 2 FWHM acceptance)
with up to 200 W on the mirror

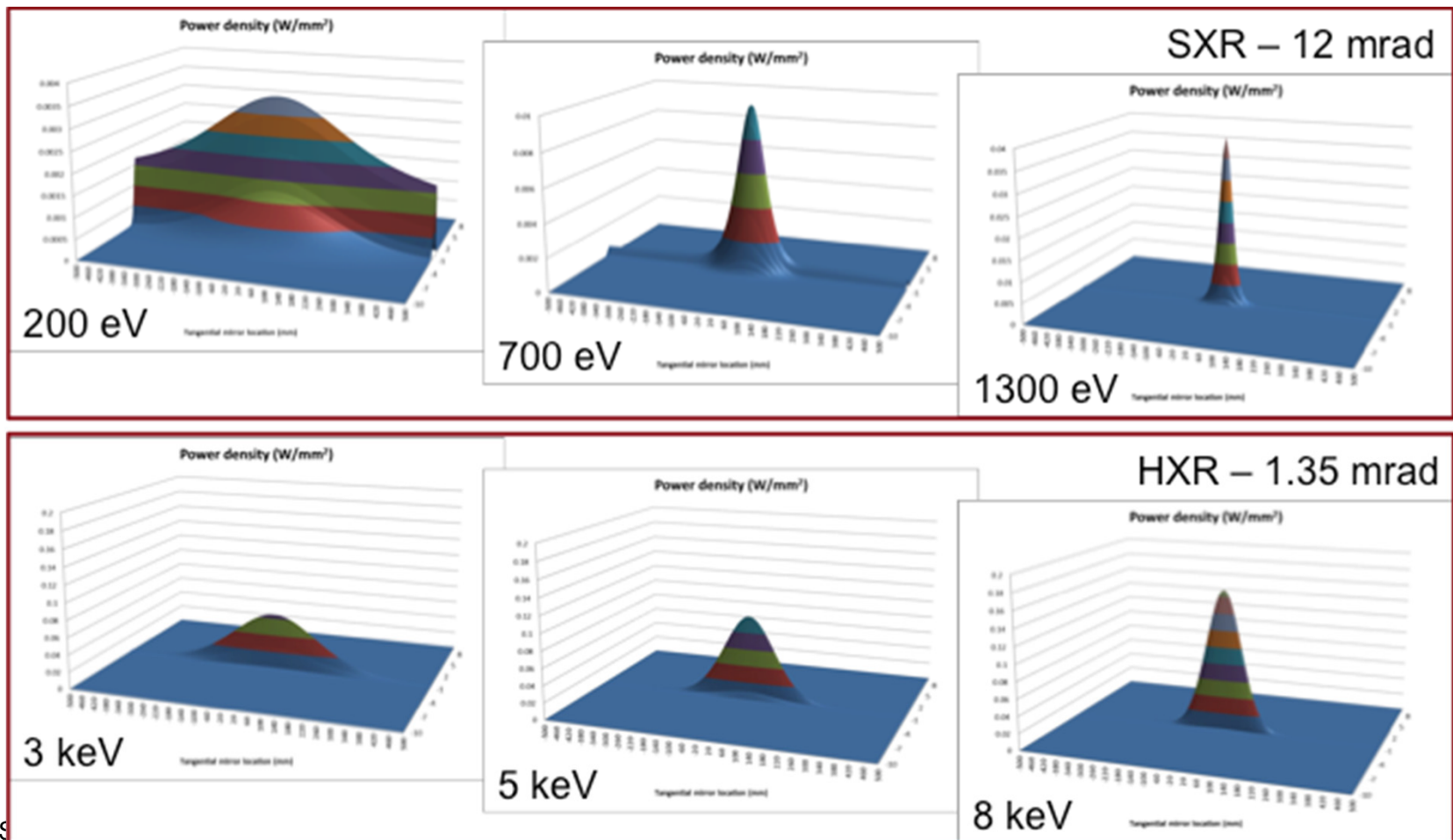


Performance to be maintained with the heat load



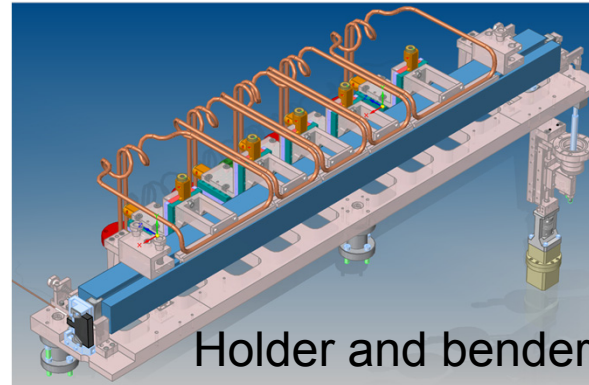
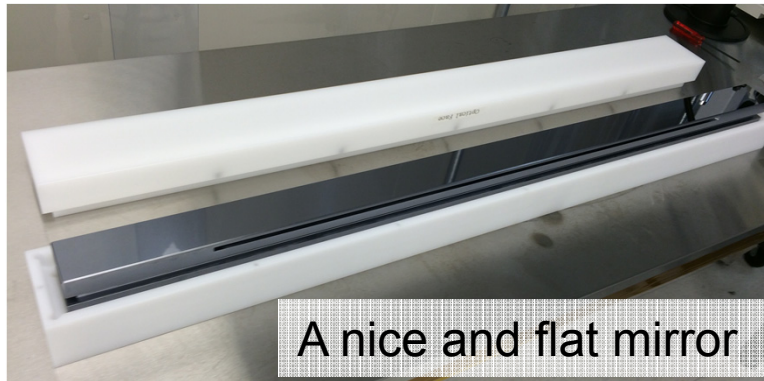
Absorbed power profiles – Flat mirror case

Wide energy range leads to a highly variable beam footprint (from 70 to 400 mm FWHM)

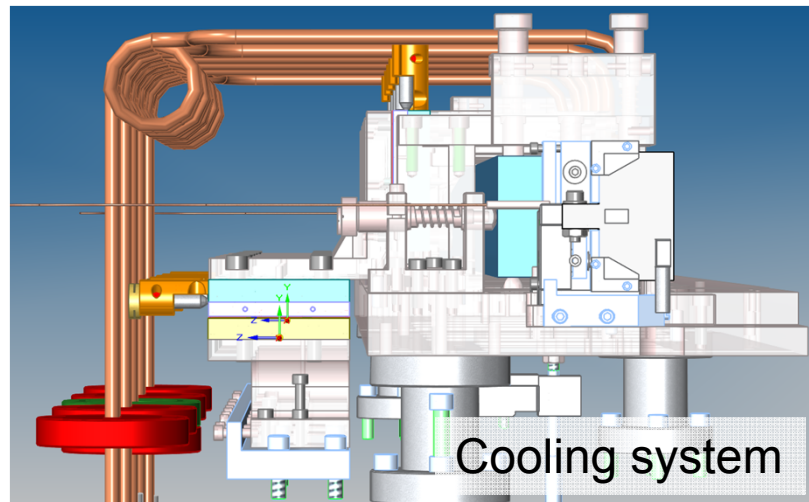


Introduction to the Flat mirror system

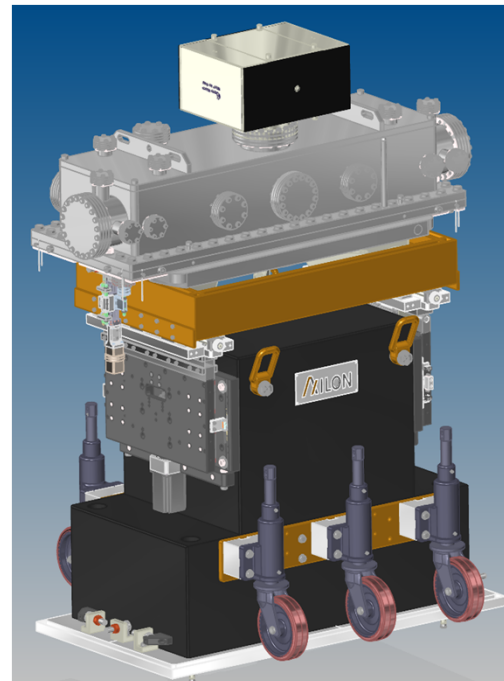
It is composed by:



Design, assembled and tested in house (an entire unit has been realized and tested)

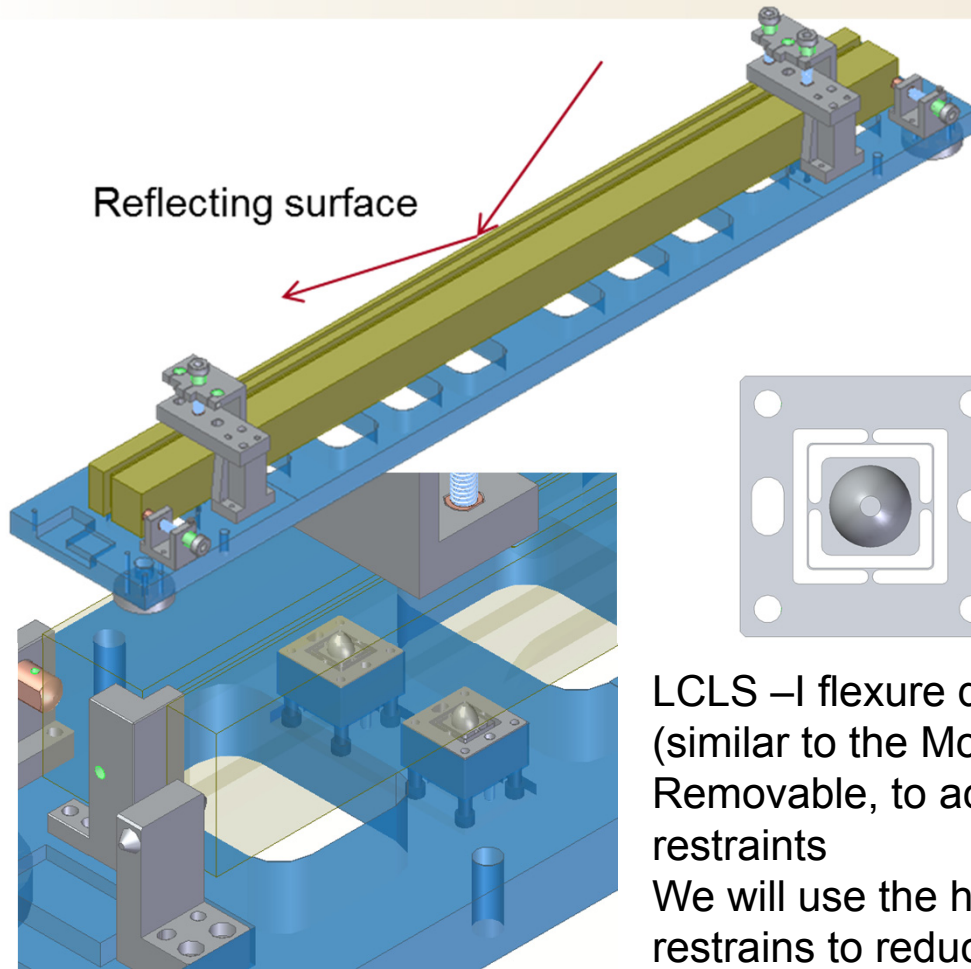


Designed, assembled and tests in house (an single unit has been realized)

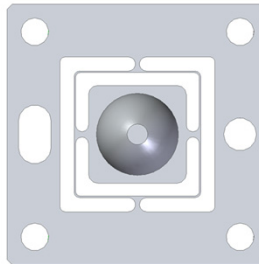


A robust and stable vacuum and mover system (identical system has been procured for the HOMS upgrade project of LCLS) Build to specs.

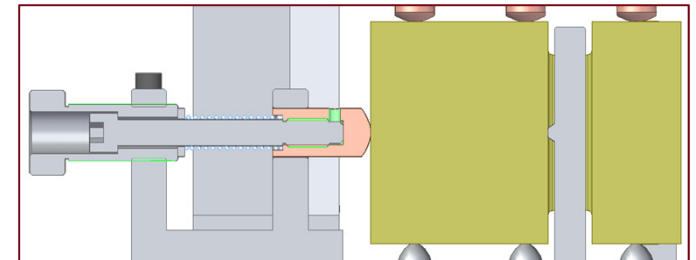
Flat Mirror - Holder



3 Vertical and 2 horizontal restraints
Positioned outside the Bessel points,
to minimize the induced deformation in
the central area



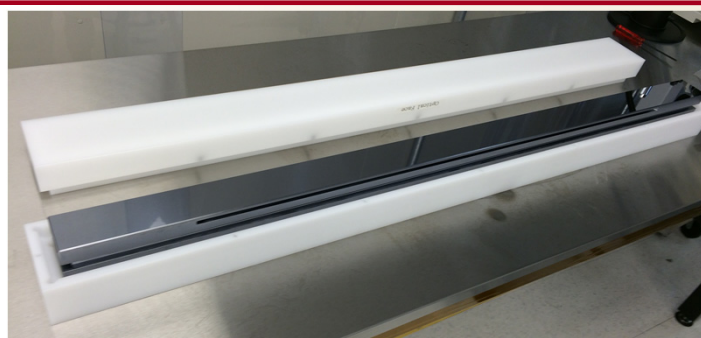
LCLS –I flexure design
(similar to the McCarville design)
Removable, to accommodate different
restraints
We will use the higher compliance
restraints to reduce deformation



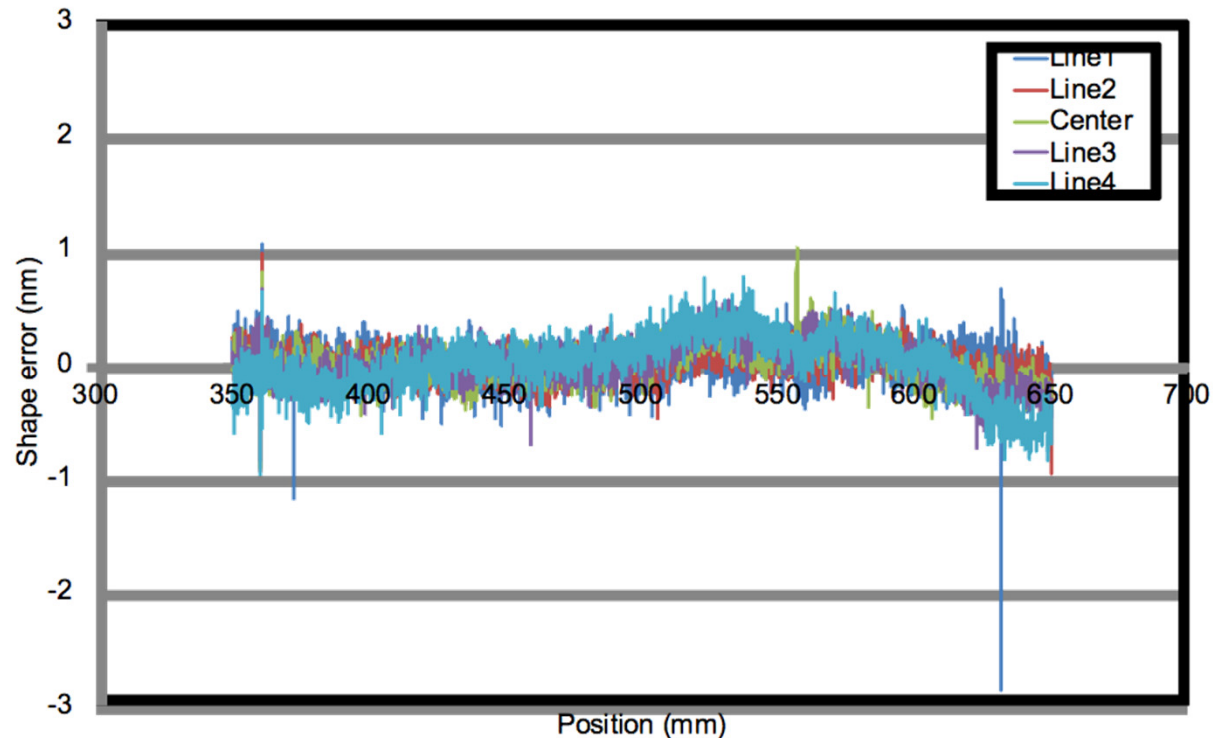
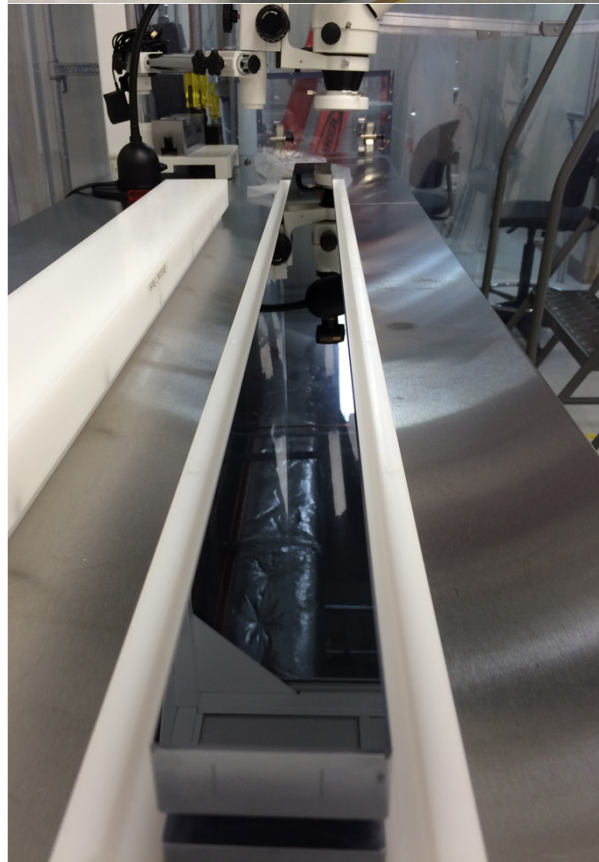
Horizontal pusher
Using the notch
around the mirror

System tested with Jtec mirrors. No appreciable deformation seen.
Careful installation procedure has to be followed

The mirrors exist!

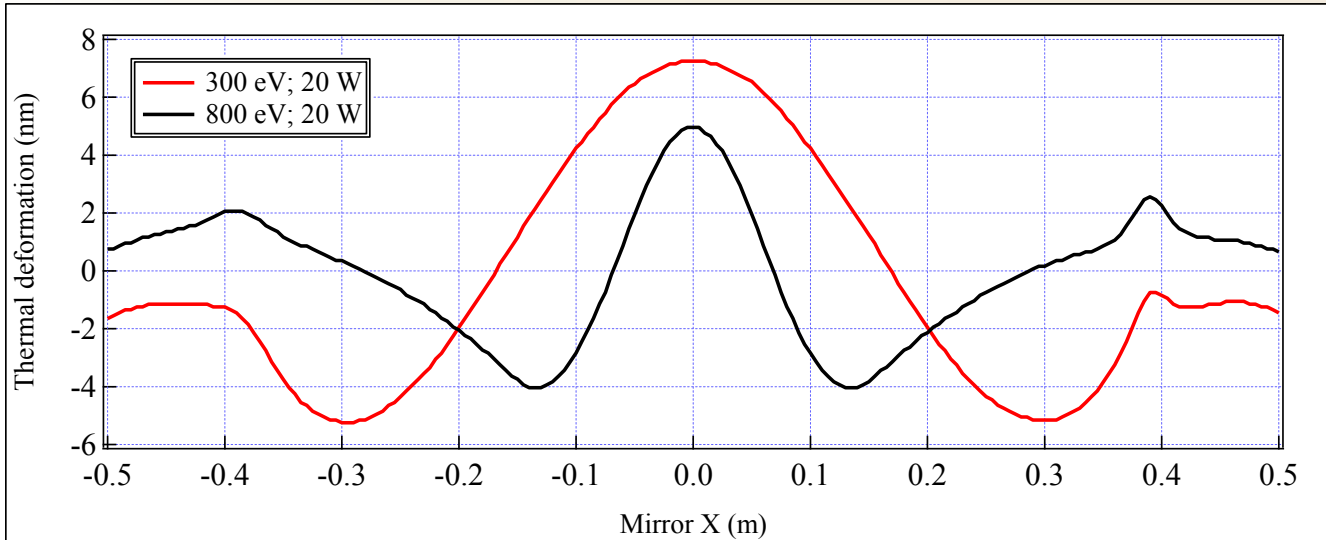


- First three (3) 1-m long mirrors for HOMS upgrade received in May 2016 at SLAC
- They are the best mirrors ever manufactured in the world!
- < 0.2 nm rms in the central 300 mm and < 0.6 overall
- $R > 900$ Km

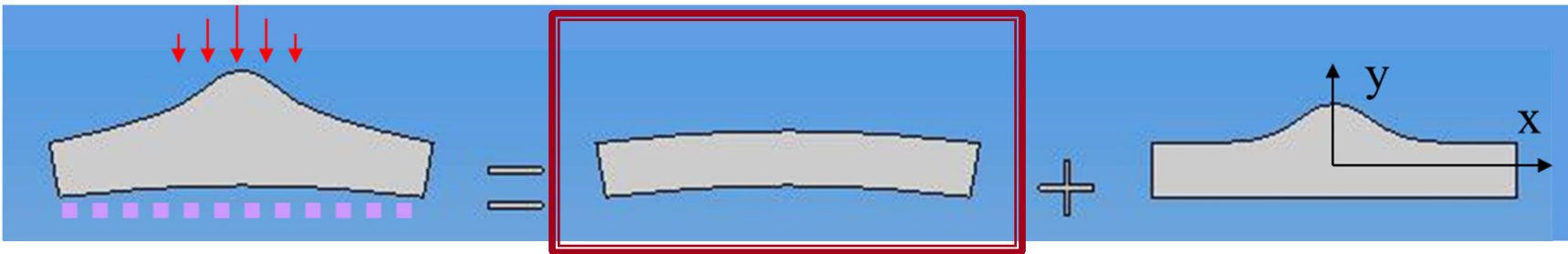


Measured shape errors at parallel lines at the vendor premise

Thermal deformation – spherical part



The deformation include a spherical thermal bump (dash lines) and residual shape errors

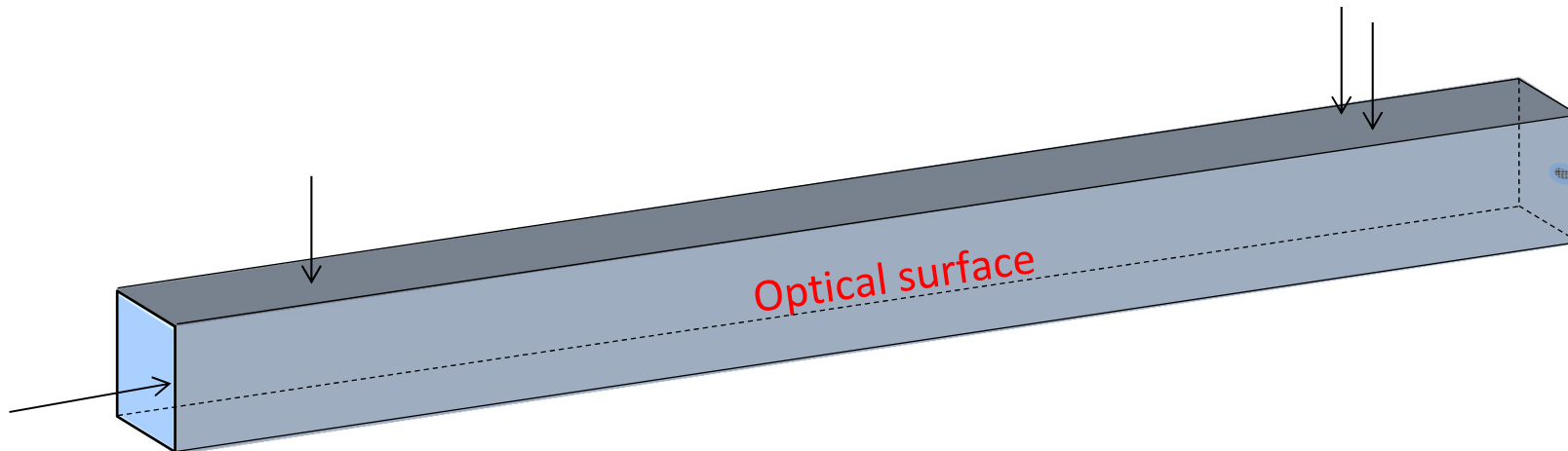
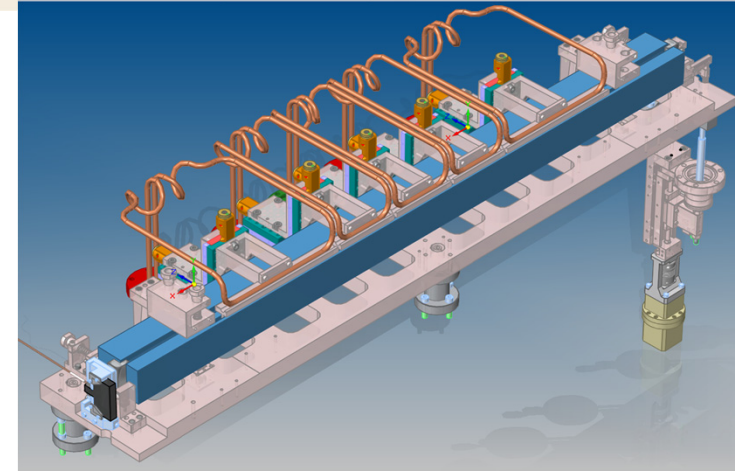


Flat Mirror - Bender

We Need to preserve the flatness – $R > 900$ Km

(with and without heat load)

- Successfully optimized installation procedure to do not introduce stress on the mirror (measured $R > 1000$ km!)
- Tested effect of each restrain (no visible mirror deformation)
- Tested the bender reproducibility (impressive preliminary results)

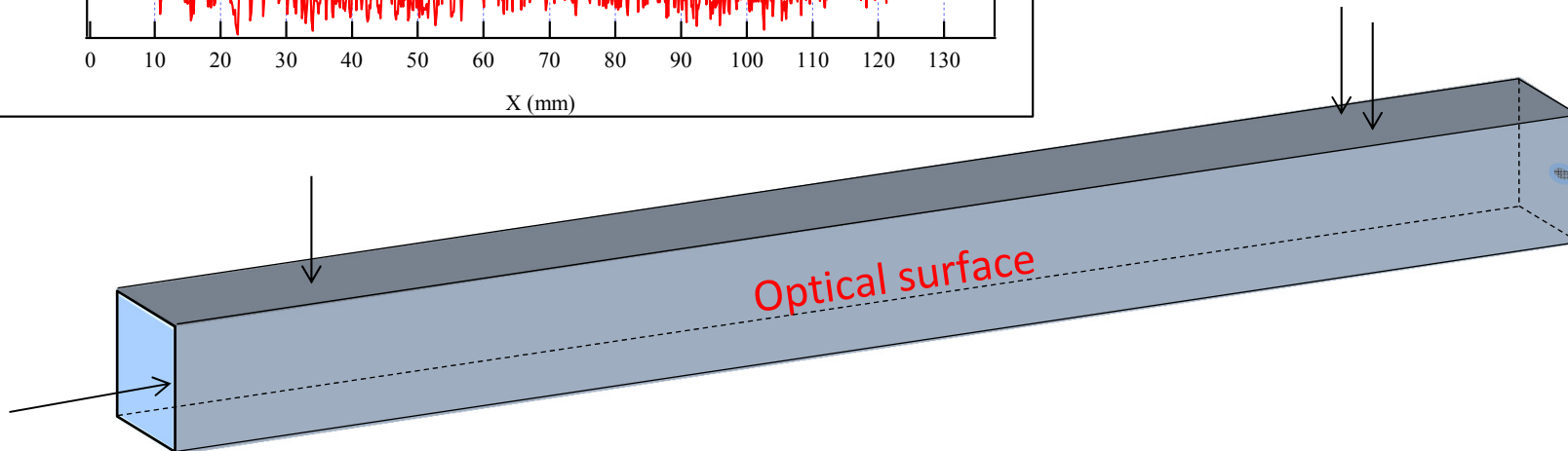
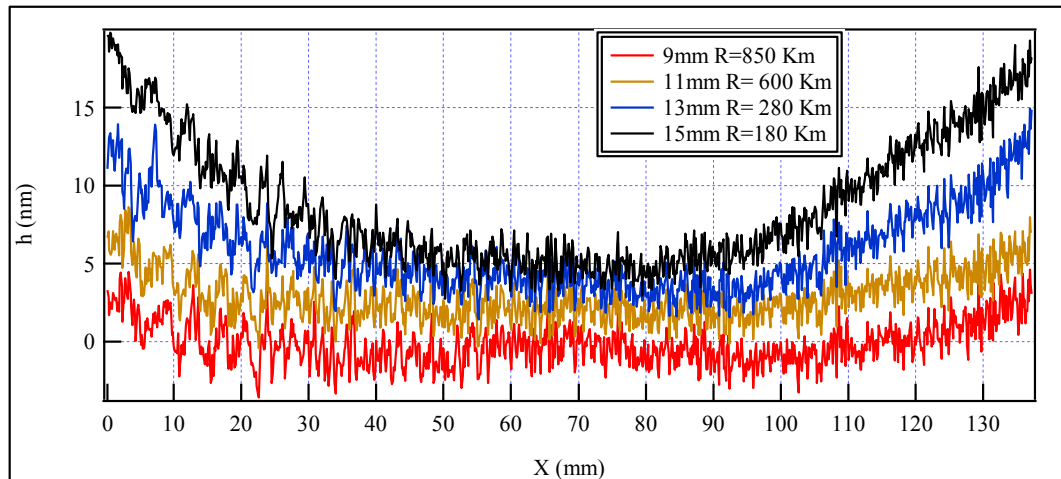
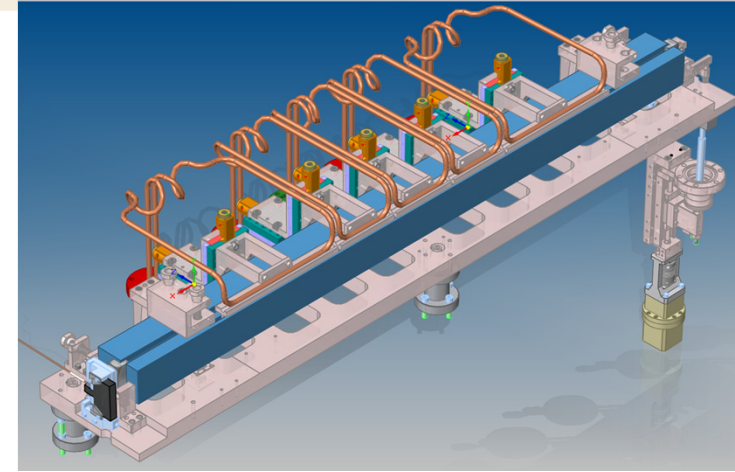


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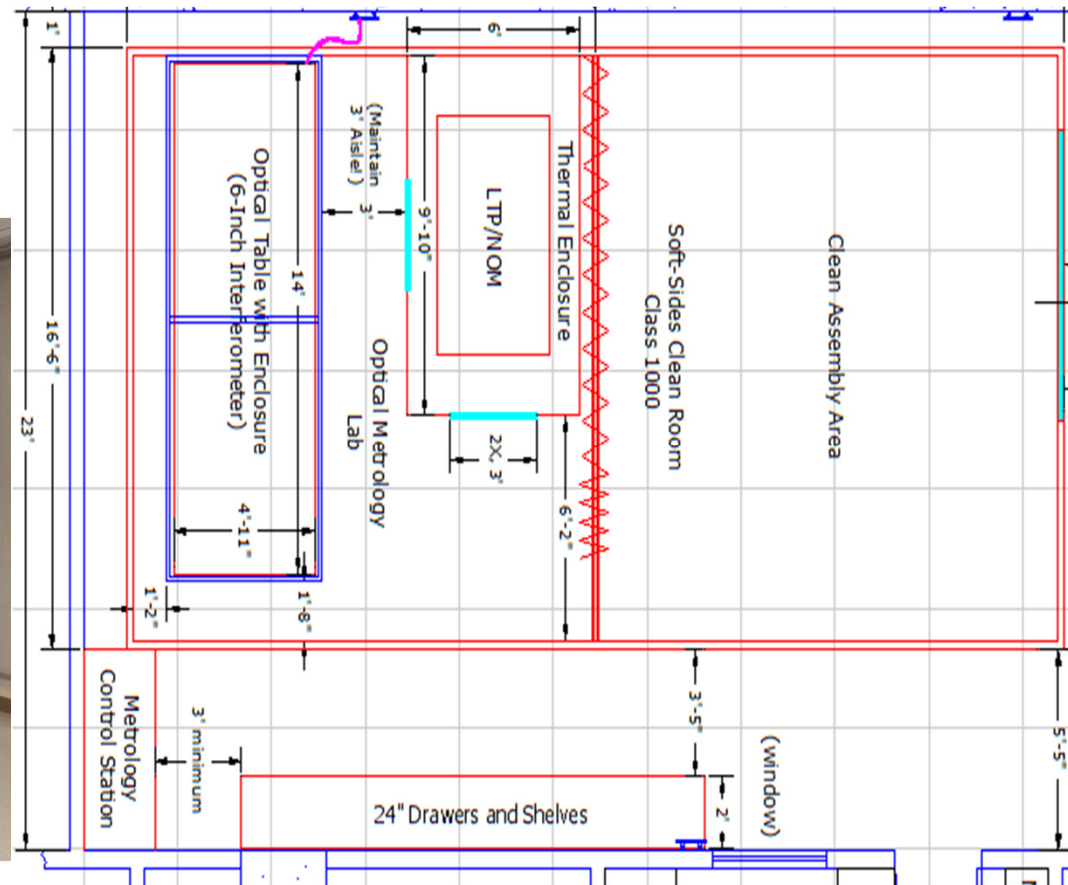


Metrology Laboratory

Class 1,000

Based on 16 HEPA filter

Enclosure provided by
clear vinyl curtain



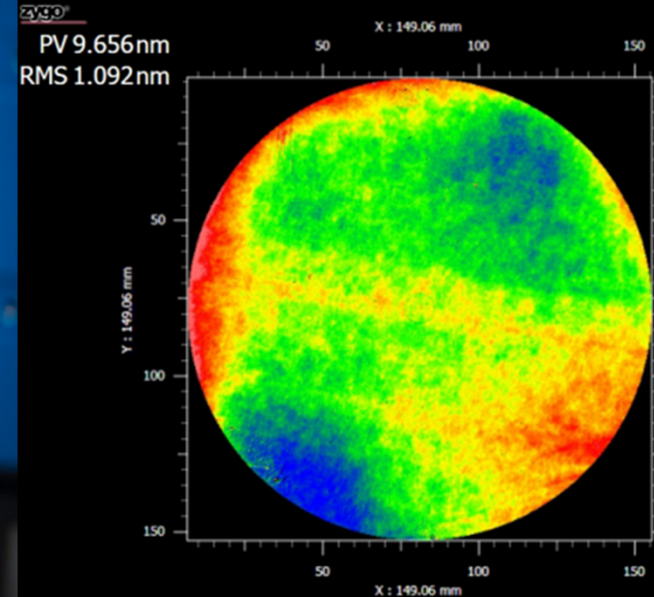
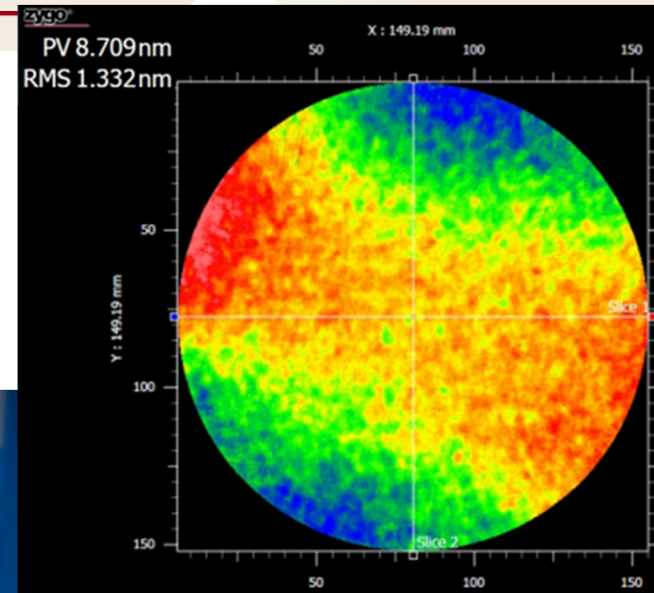
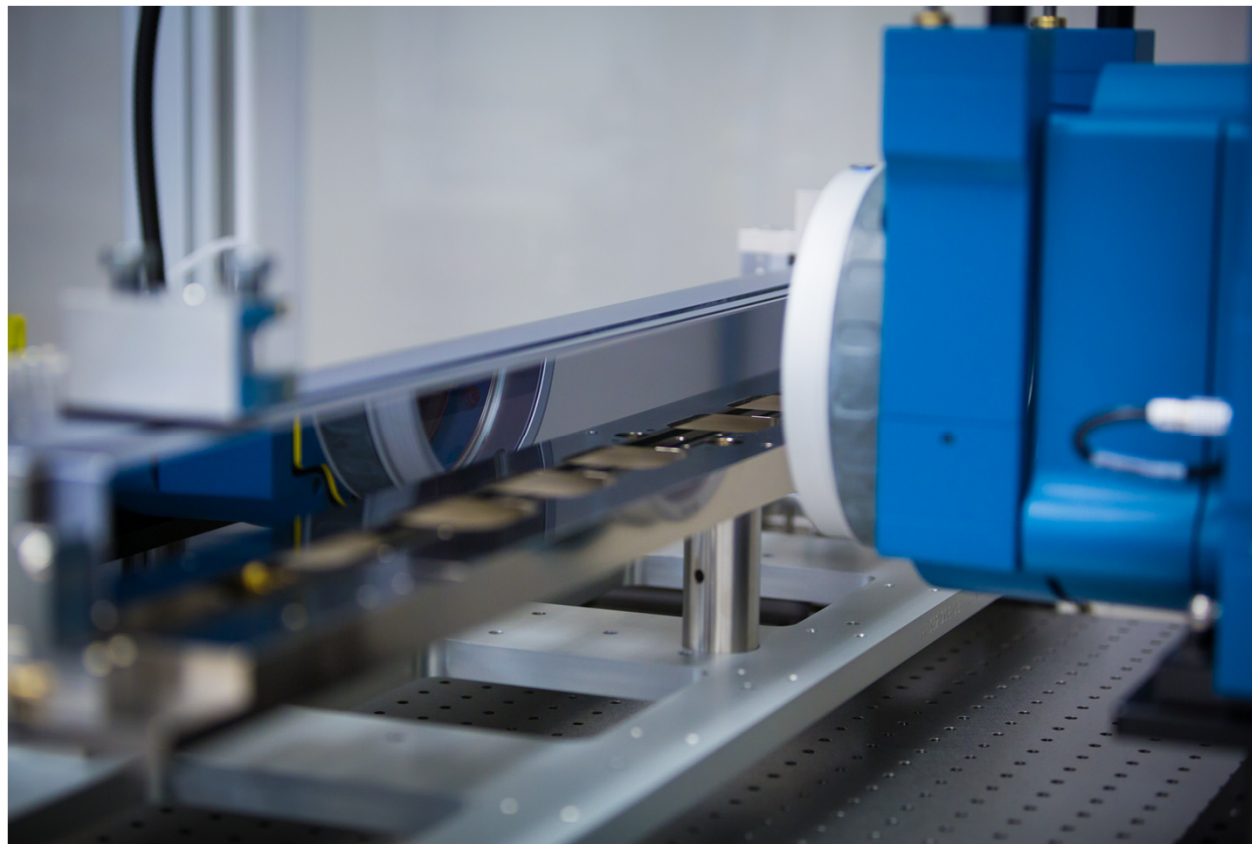
Humidity and temperature controlled by maintaining stable the circulating air.

Temperature stability: $\pm 0.1^{\circ}\text{C}$ with up to 8 people in the room (by design) at 85°F

Humidity: $\pm 2.5\%$ at 50%

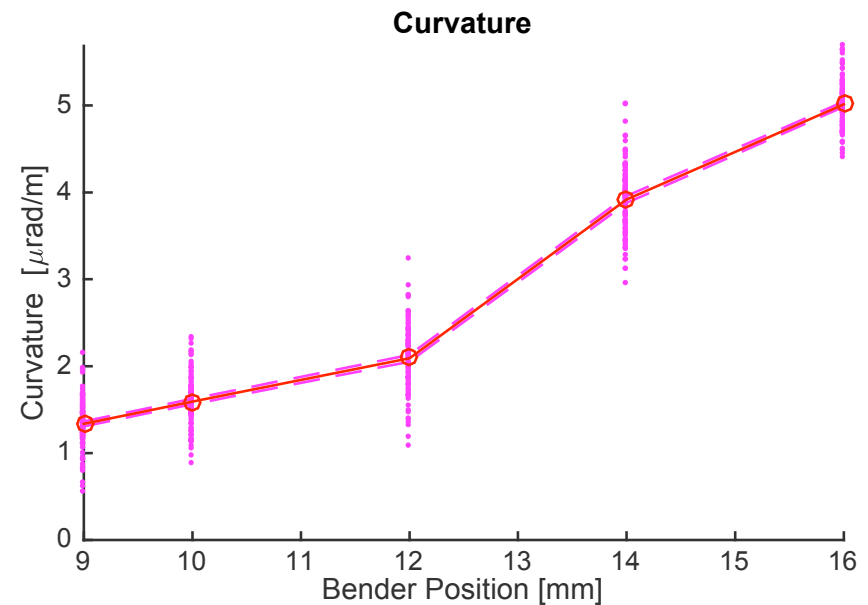
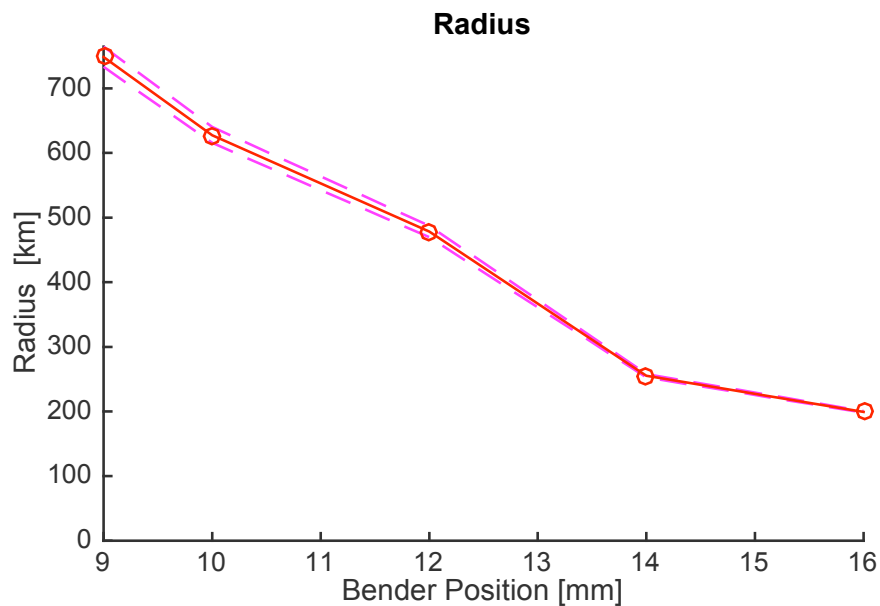
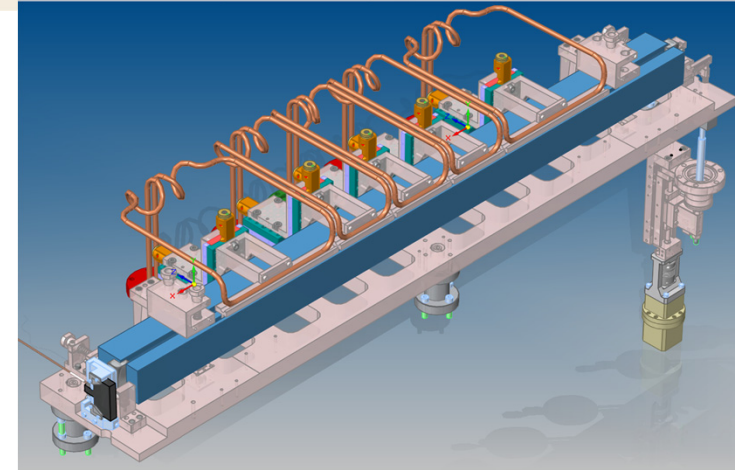
Fizeau Interferometer

- We have 3 thick, very flat references for our interferometer.
- 1 $\lambda/300$, 1 Dynaflect $\lambda/300$ (4%LT), and one $\lambda/50$.



Test results on the bender

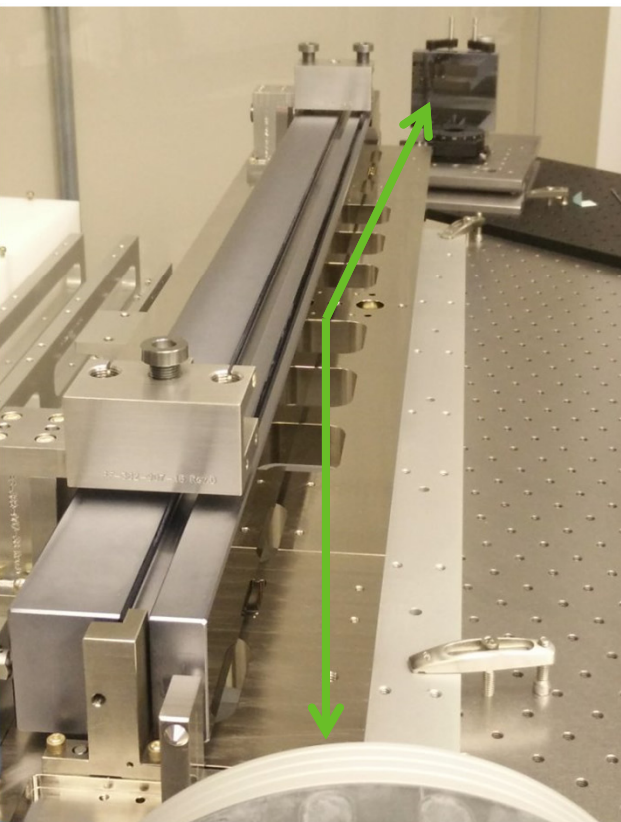
Several rounds of measurements performed
(>1000 measurements)
Designed to compensate radius of 250 Km
(expected to be able to compensate radius as
short as 150 Km)



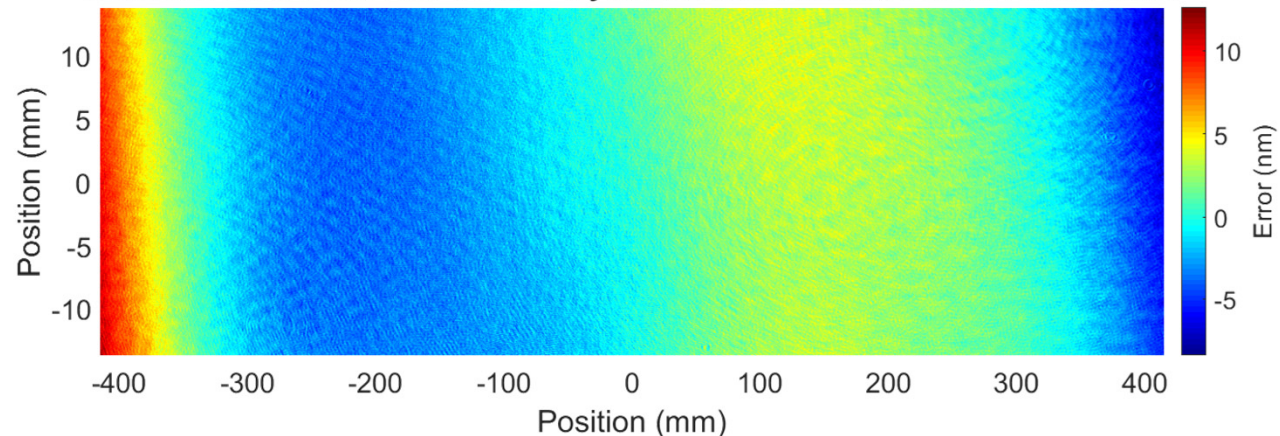
Grazing Incidence Measurements

The bender induces a small, low spatial frequency residual cubic error term measured in grazing incidence ($\sim 10\text{nm PV}$)

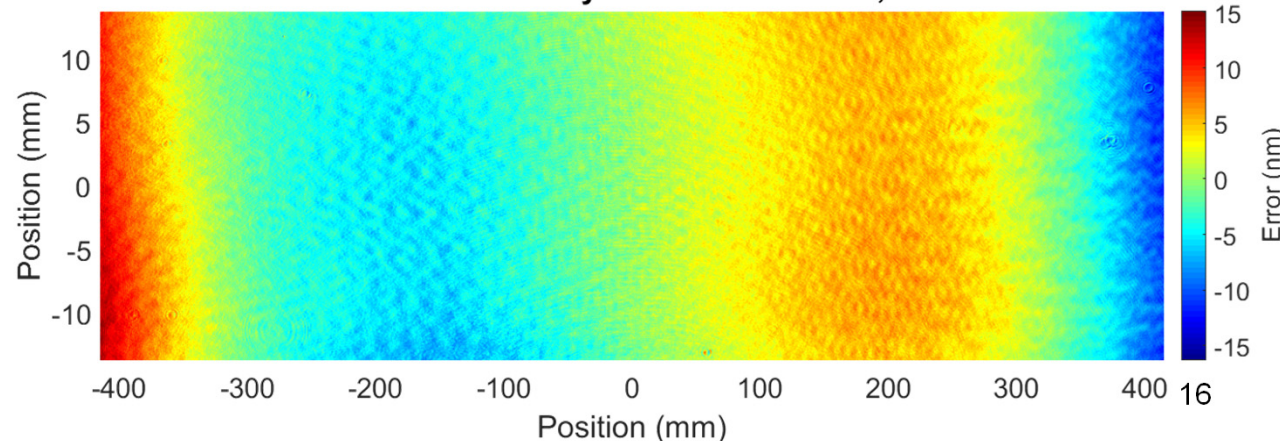
Resolution is $\sim 1/5^{\text{th}}$ Normal Incidence at a Grazing Angle of 5.6 degrees ($\sim 0.5\text{nm}$)



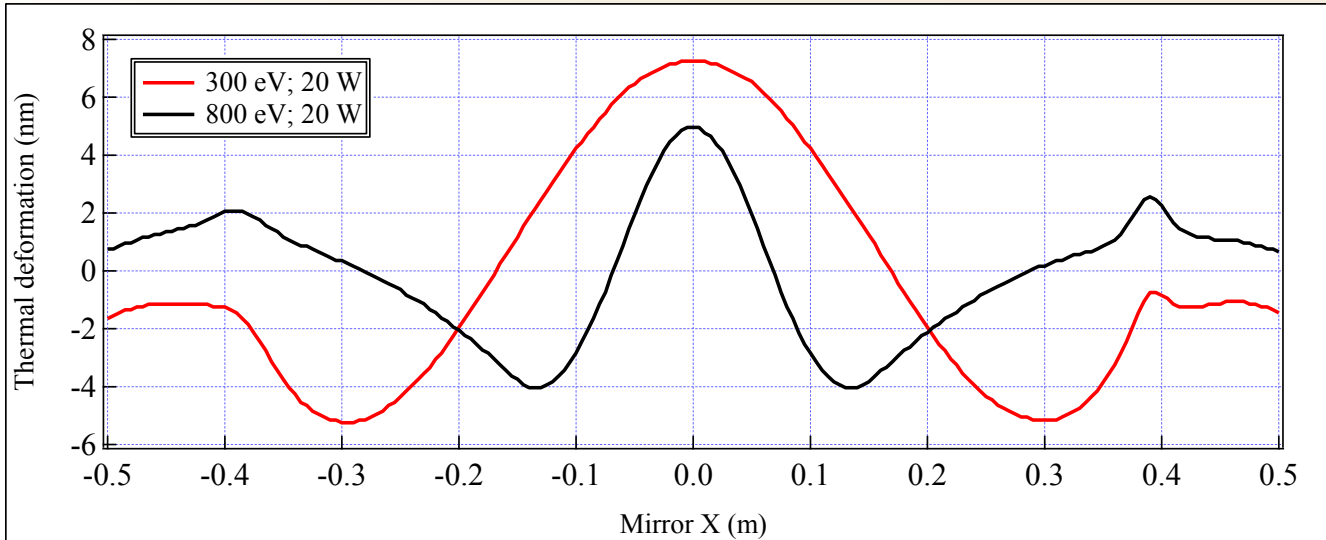
Residual Error After 2nd Order Polynomial Subtraction, 350km radius bend



Residual Error After 2nd Order Polynomial Subtraction, 160km Radius Bend



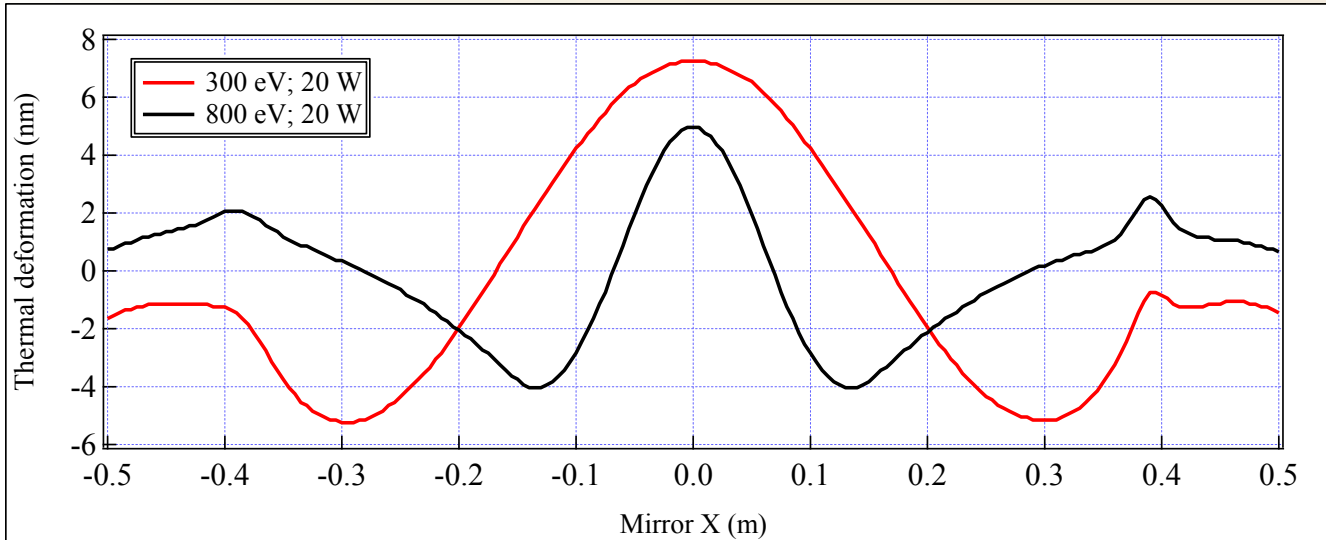
Thermal deformation – aspherical part



Photon Energy (eV)	Power (W)	Radius (km)
300	20	1,800
800	20	920
200	200	~470
1300	200	~200



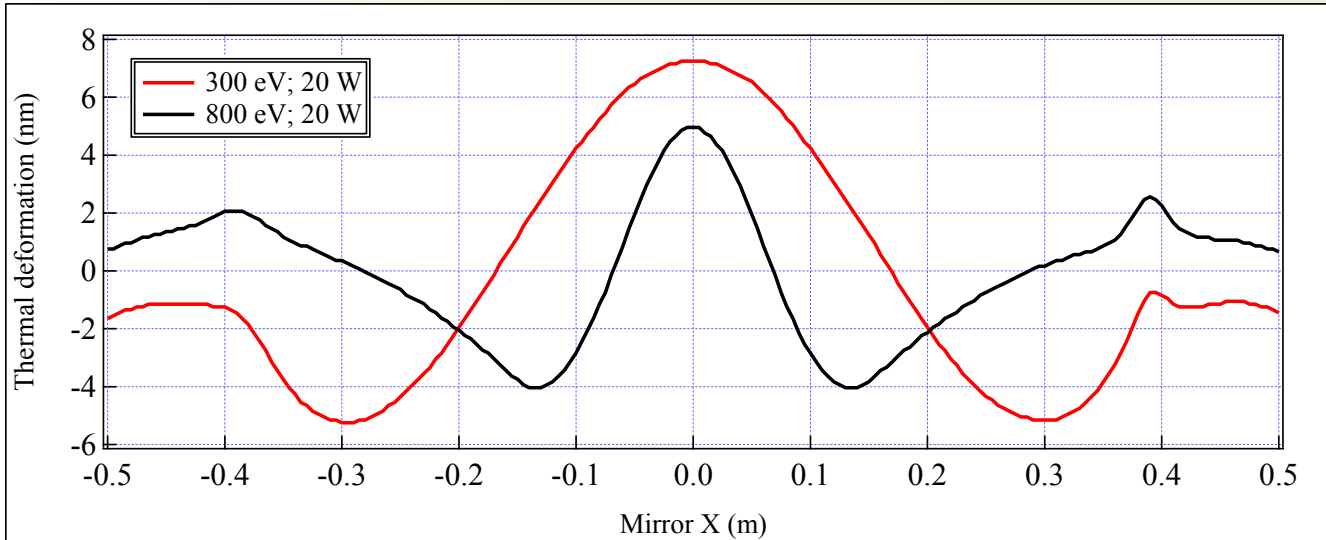
Thermal deformation



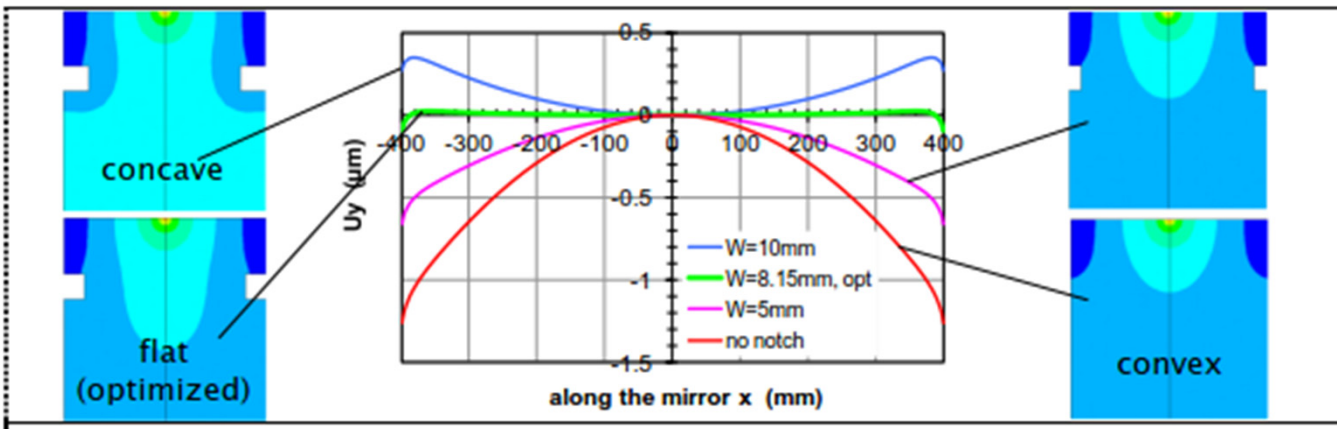
The deformation include a spherical thermal bump (dash lines) and a residual shape errors



Thermal deformation

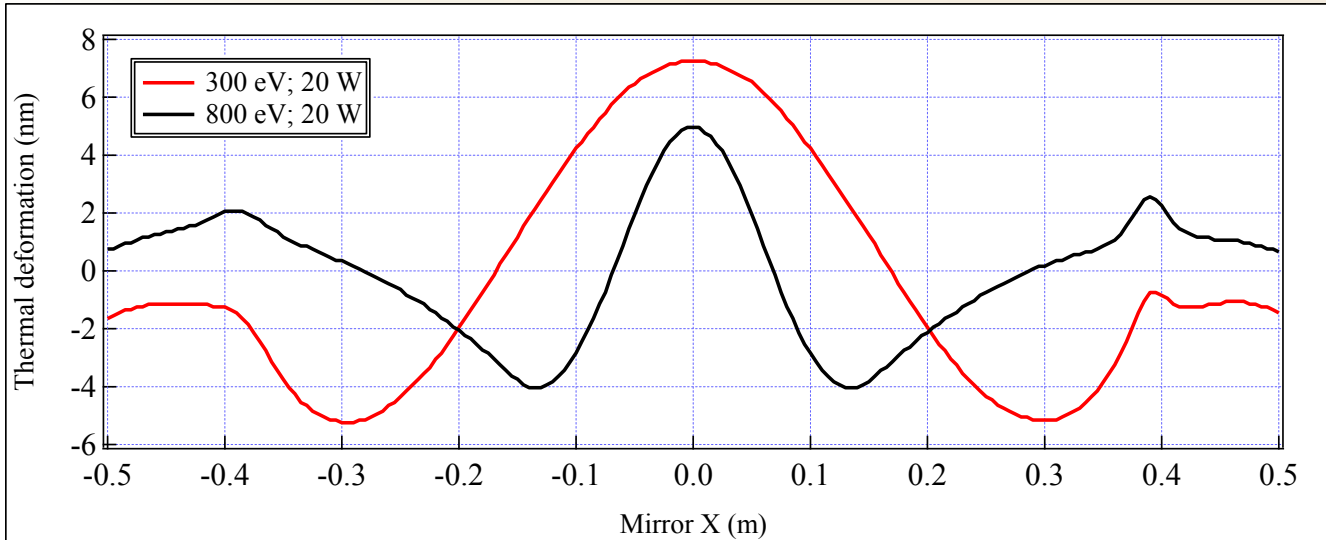


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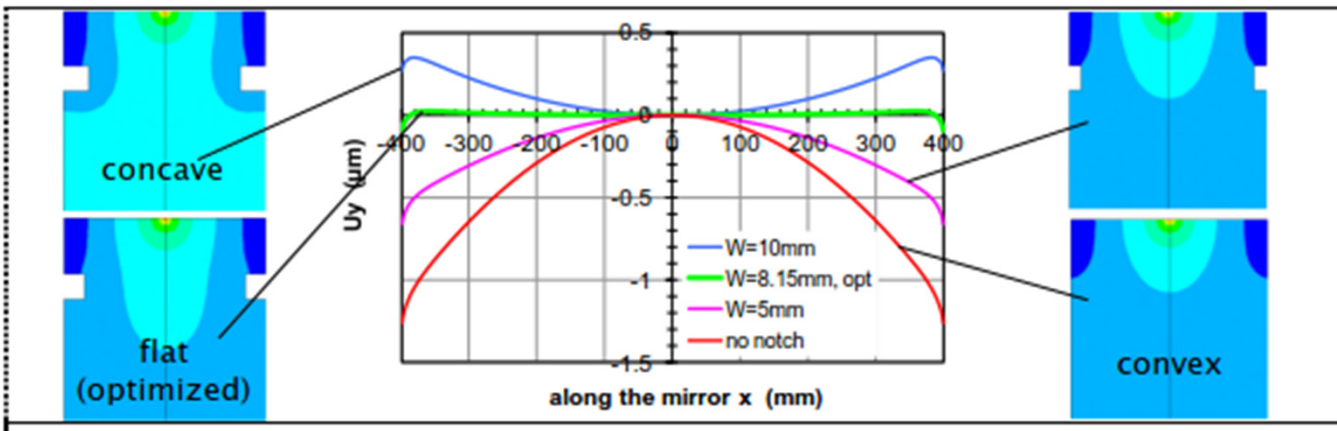


Thanks to the notch, the induced spherical thermal bump can be minimized for a range of beam footprints

Thermal deformation



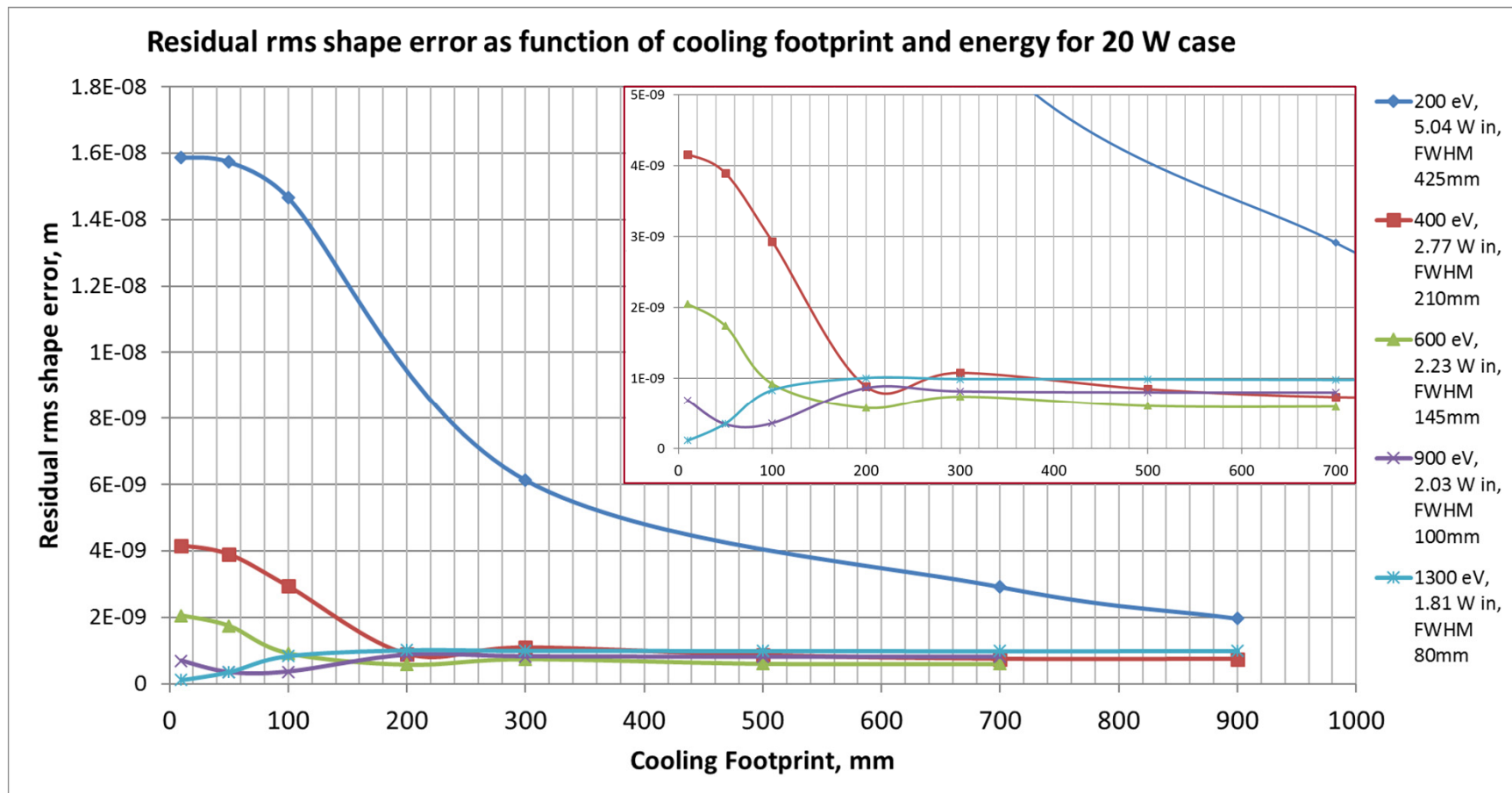
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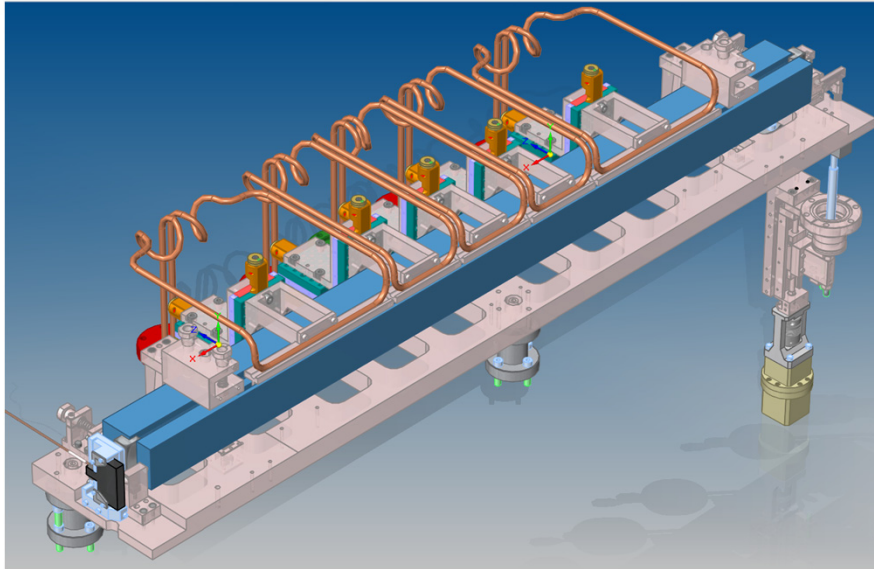
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Effect of Cooling Footprint on Performance

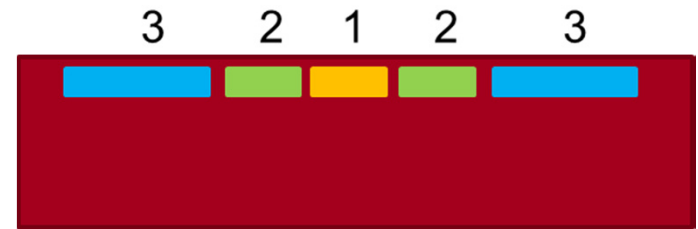
The ideal scenario is when the cooling length matches the beam FWHM.



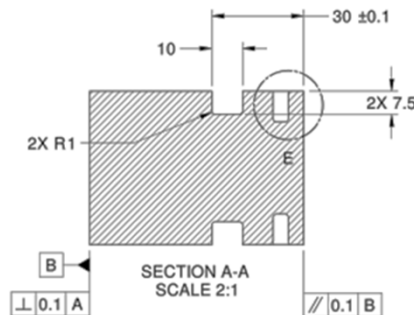
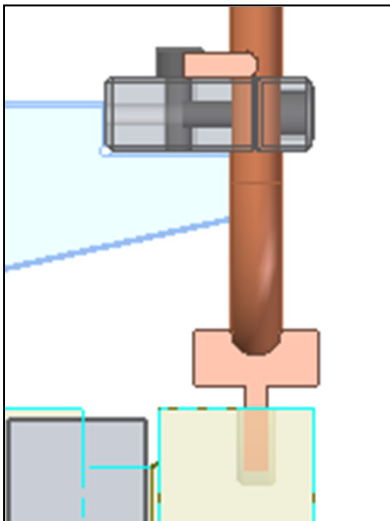
Cooling scheme



A simplified solution uses five cooling circuits (3 cooling lengths) to mimic the beam footprint

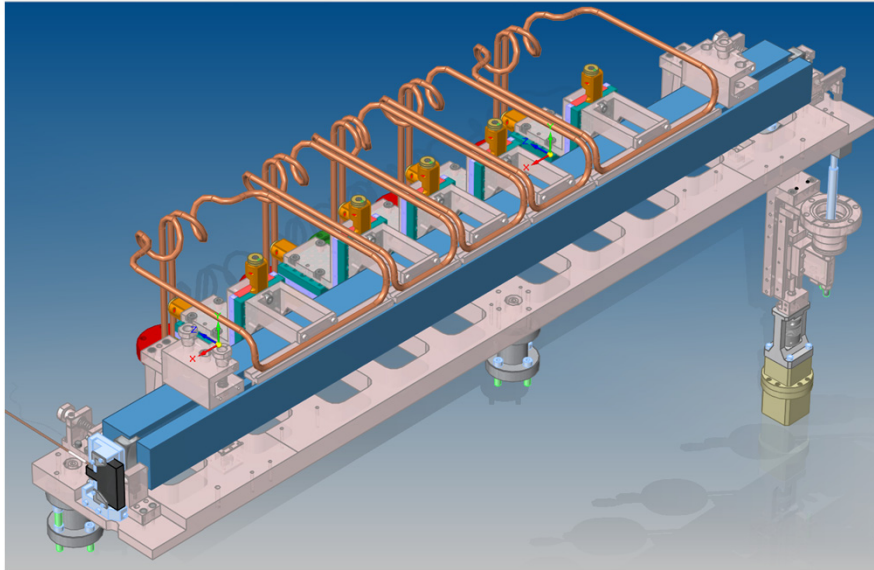


Energy Range (eV)	Cooling Footprint (mm)
< 400	700
~ 400 – ~ 800	300
> 700	100



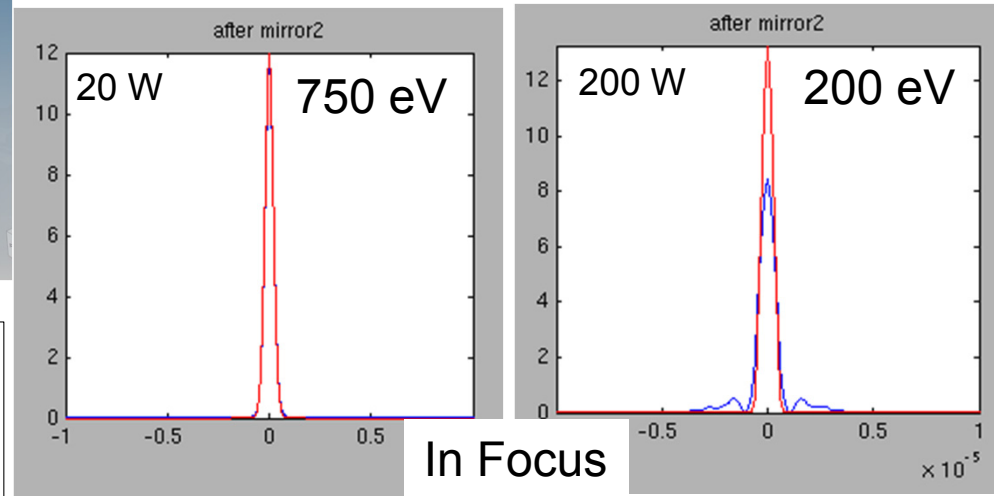
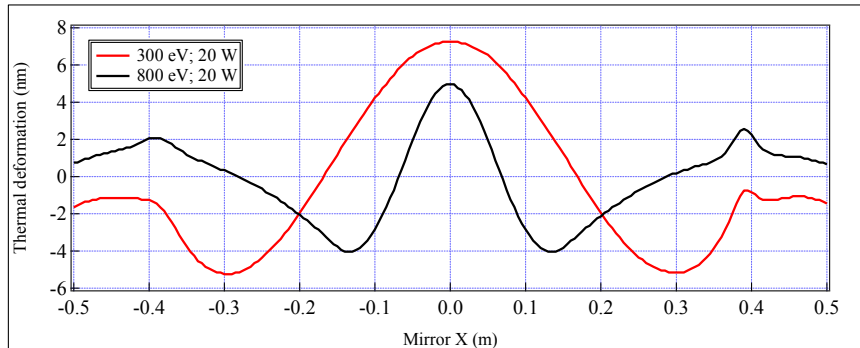
The mirror is cooled from one side only. It has a trough located 5 mm from the surface hosting the Galn eutectic. A notch, 20 mm from the surface, is used to reduce the thermal bump, and hold the mirror

Flat Mirrors (HOMS/SOMS) design – in focus

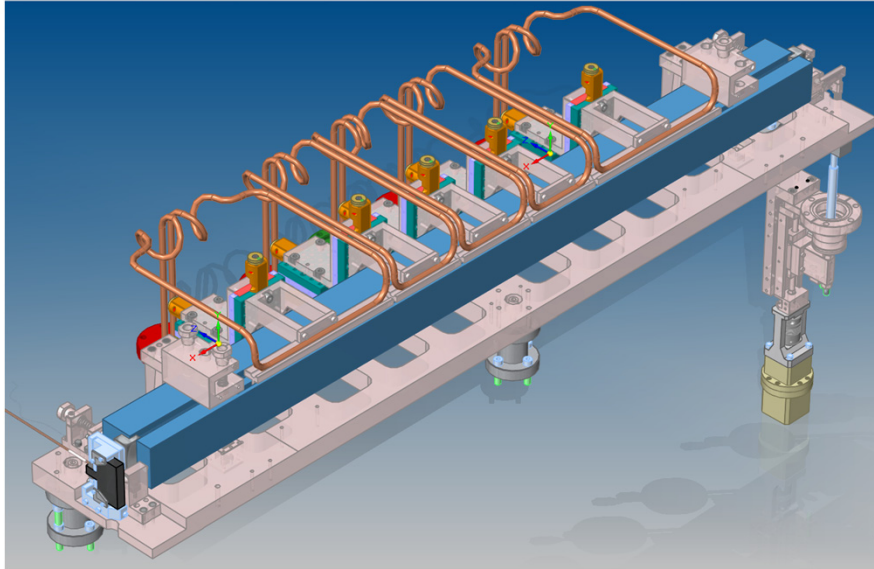


The ideal scenario is when the cooling length matches the beam FWHM.
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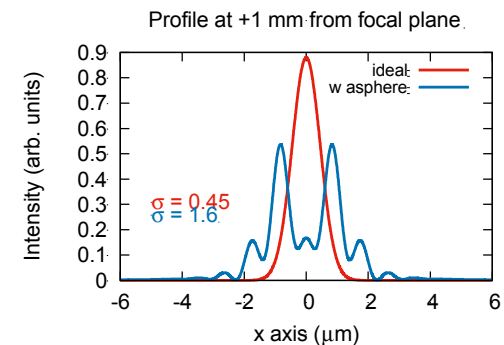
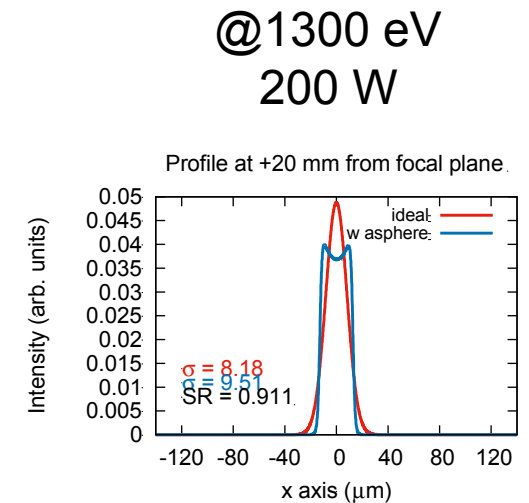
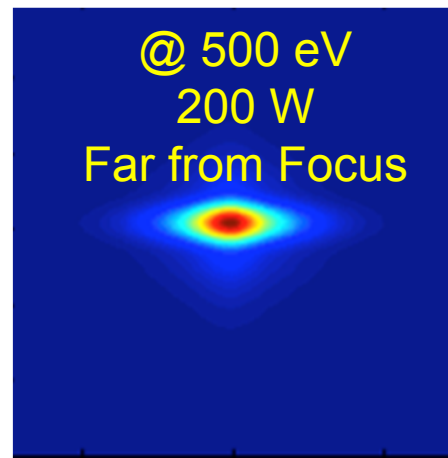
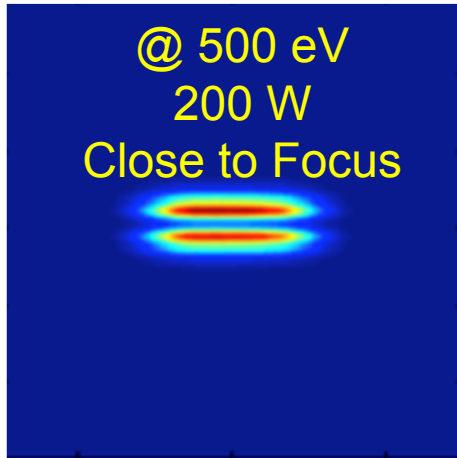
Red=perfect mirrors; Blue=with thermal deformation



Flat Mirrors (HOMS/SOMS) design - out of focus

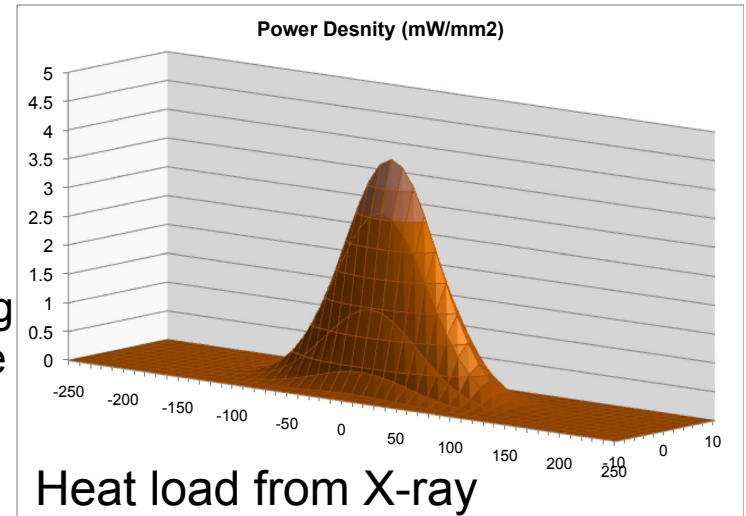
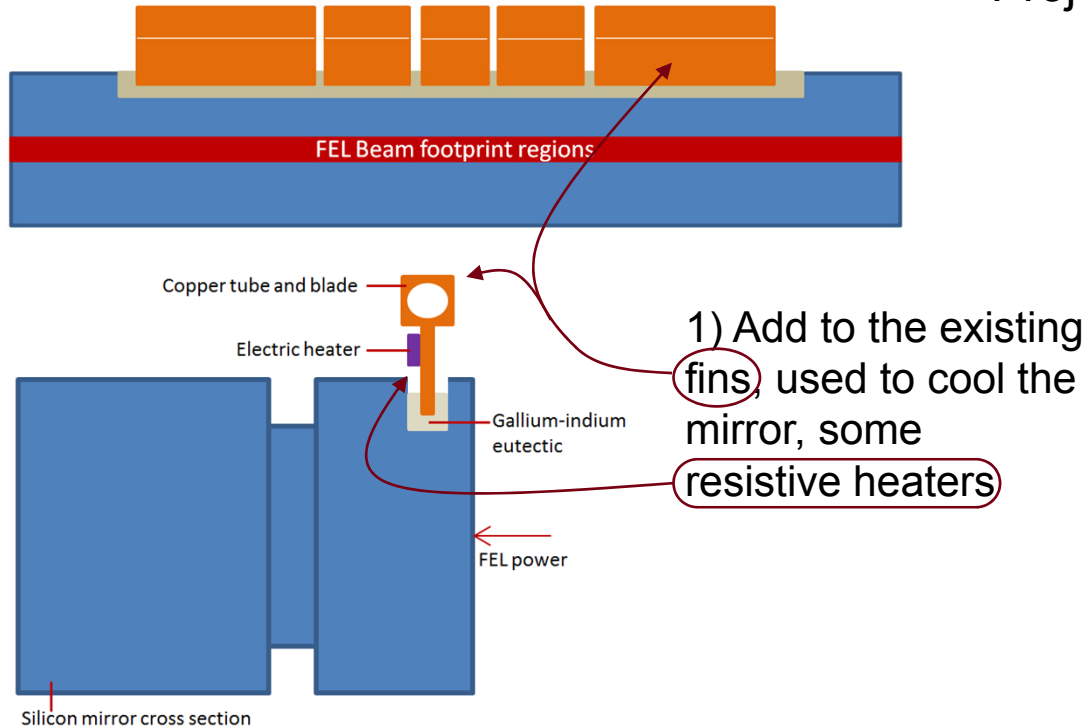


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A simplified solution uses five cooling circuits (3 cooling lengths) to mimic the beam footprint



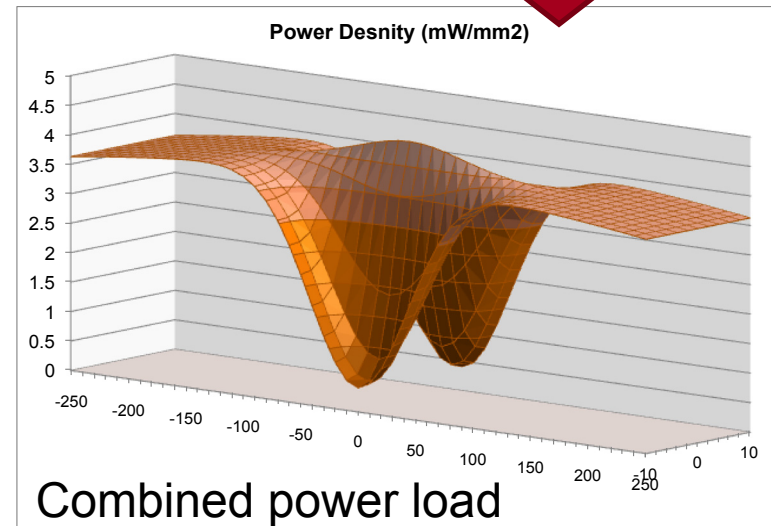
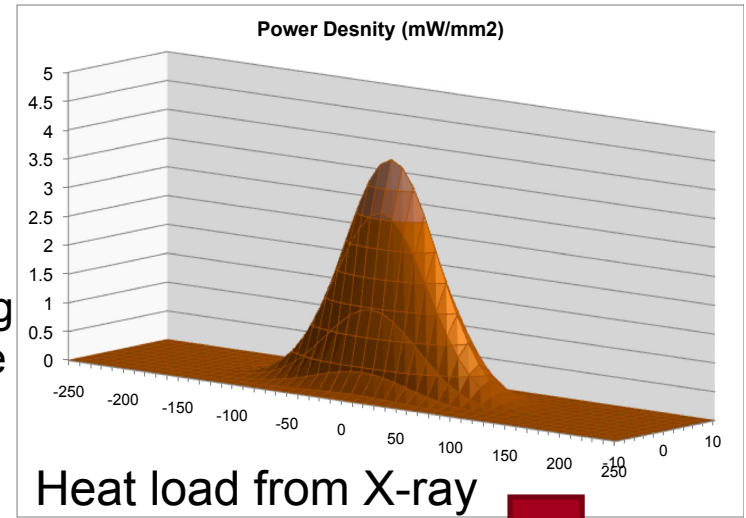
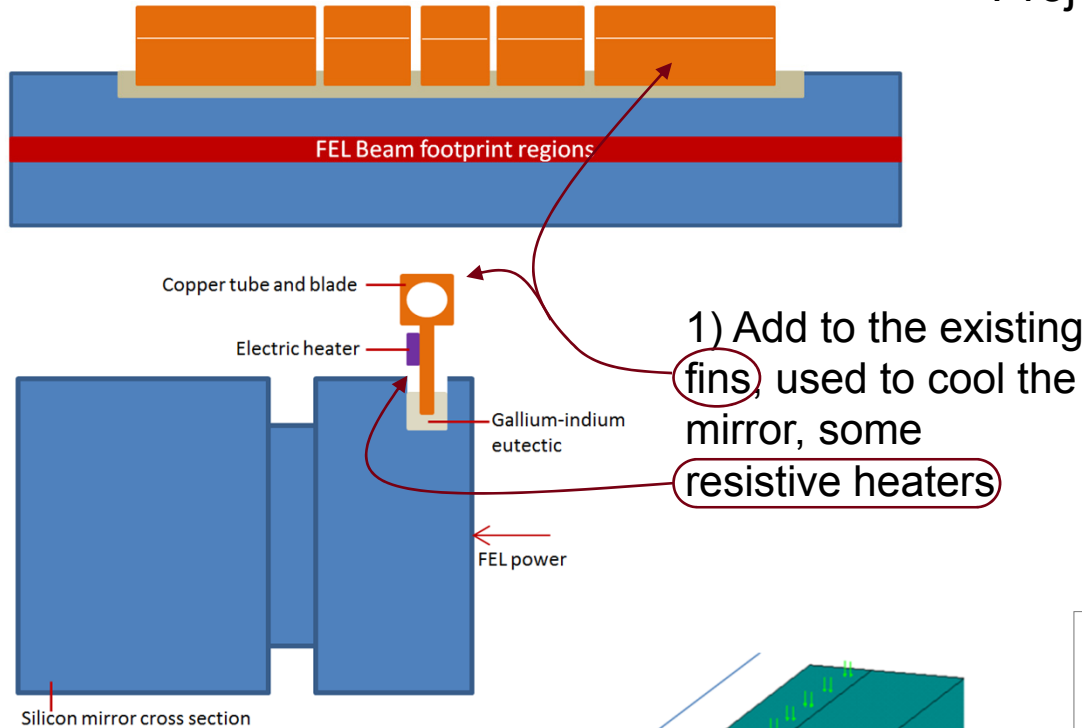
Future development: REAL (Resistive Element Adjustable Length) Cooled Optics

Project funded by BES over two years

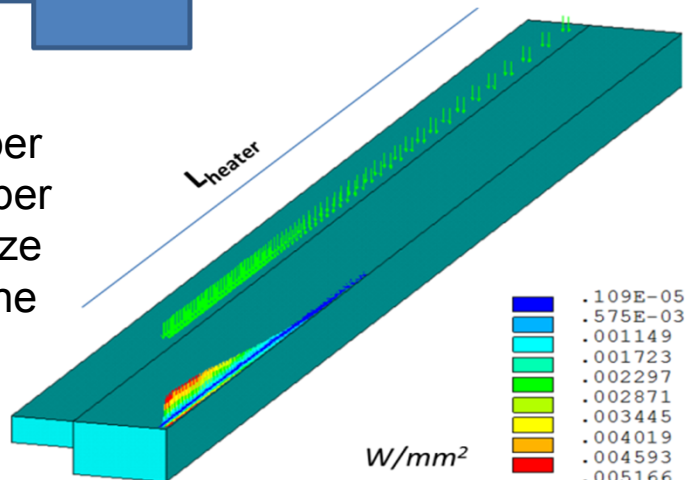


Future development: REAL (Resistive Element Adjustable Length) Cooled Optics

Project funded by BES over two years

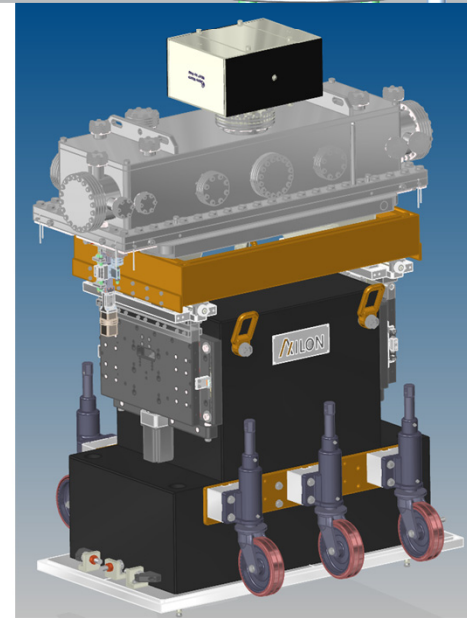
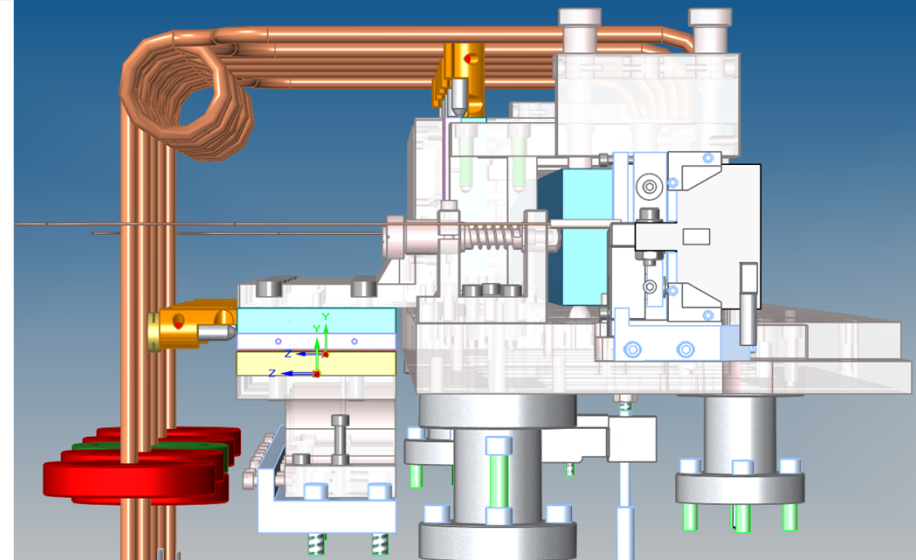


2) Apply the proper power to the proper heaters to equalize deformation on the mirror



Next steps

- Test prototype mirror with Gallium indium
 - Measure static vibration performance
 - Measure influence of cooling on vibration
 - Ensure proper chiller performance
- Test REAL cooling technique
- Install mirrors in LCLS I with no cooling in March 2017
- Full LCLS II upgrade with cooling beginning ~June 2018



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Questions?